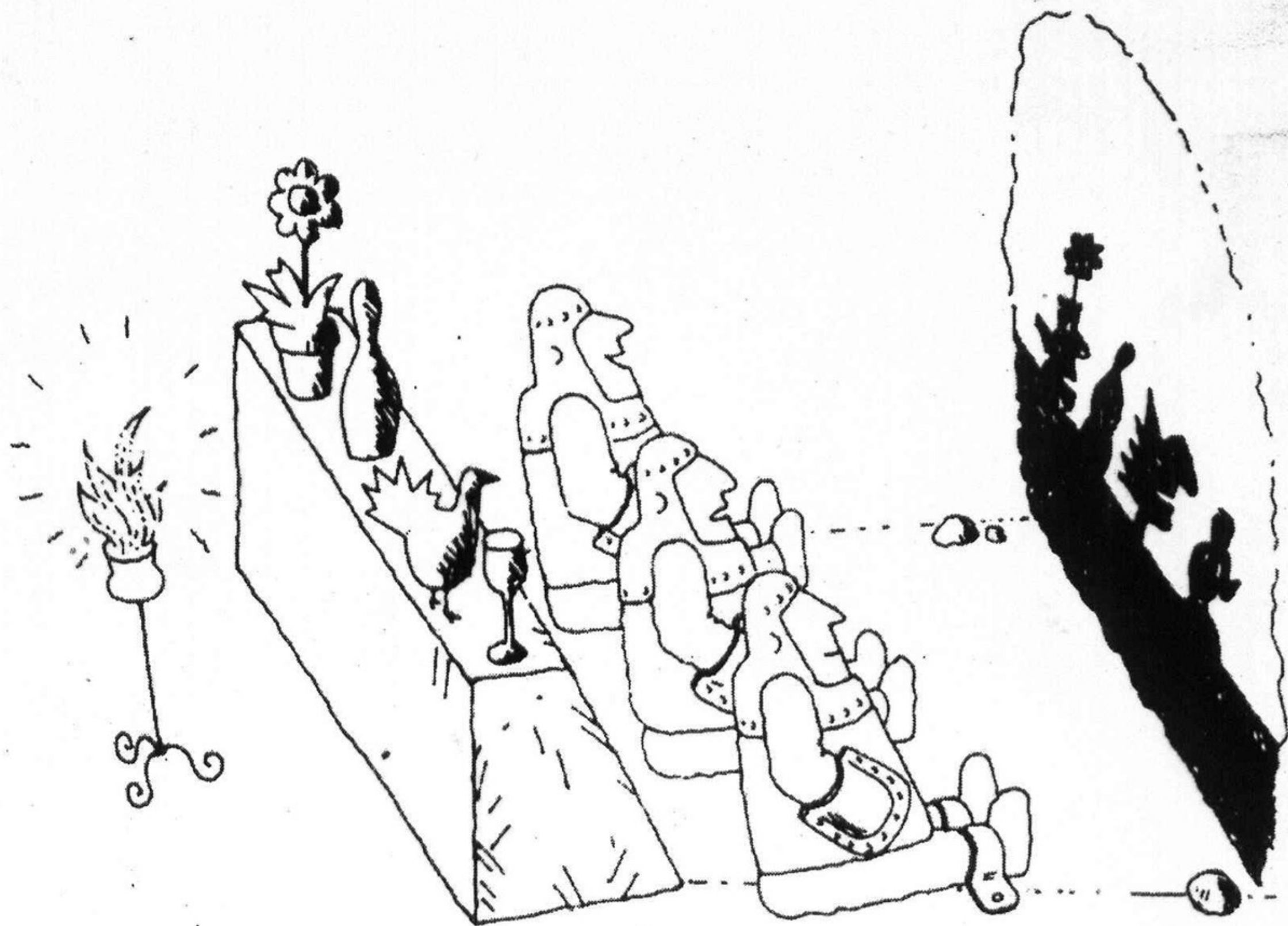




Data analysis and Statistics in paleoclimate

Jessica E. Tierney



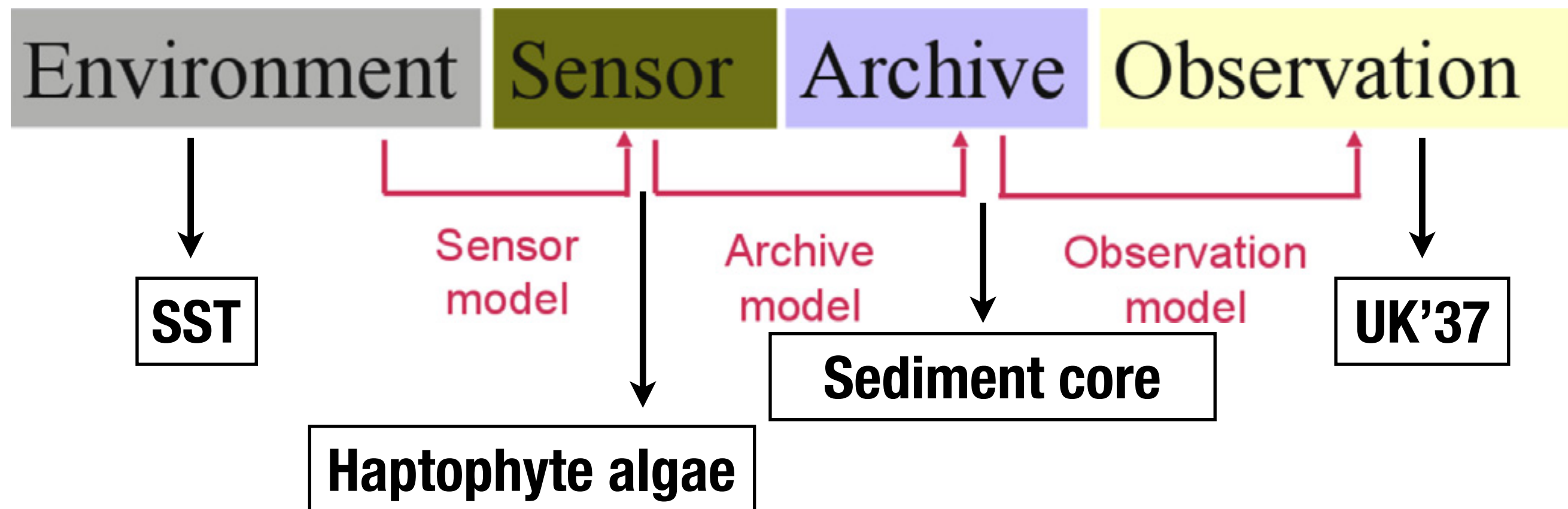


Arc of this lecture

- **The “proxy system”: Understanding the nature of the signal.**
- **Bayesian statistical approaches.**
- **Reduced space methods.**
- **Data assimilation.**

Understanding the nature of the signal: “The proxy system”

Evans et al., 2013, QSR



The World of Bayes

$$P(A|B) = \frac{P(B|A)P(A)}{P(B)}$$

Bayesian vs. Frequentist

Frequentist: data are a series of events, with a frequency, and parameters are set. The probability of occurrence is immutable.

Bayesian: data are fixed and parameters must be estimated. Probability is mutable and changes with the addition of information.

Why go Bayesian?



1) Allows one to model the proxy behavior and uncertainties in an etiologically correct way:

$$\text{TEX}_{86} = \alpha_s + \beta_s * \text{SST} + \epsilon$$

2) Then we can use Bayes' rule to predict SST from TEX_{86} :

$$P(\text{SST} | \text{TEX}_{86}) \propto P(\text{TEX}_{86} | \text{SST}) * P(\text{SST})$$

3) $P(\text{SST})$ is the prior. It can be relatively uninformative if the model is robust.

4) From this process, we get meaningful uncertainty propagation, plus ensembles of results.

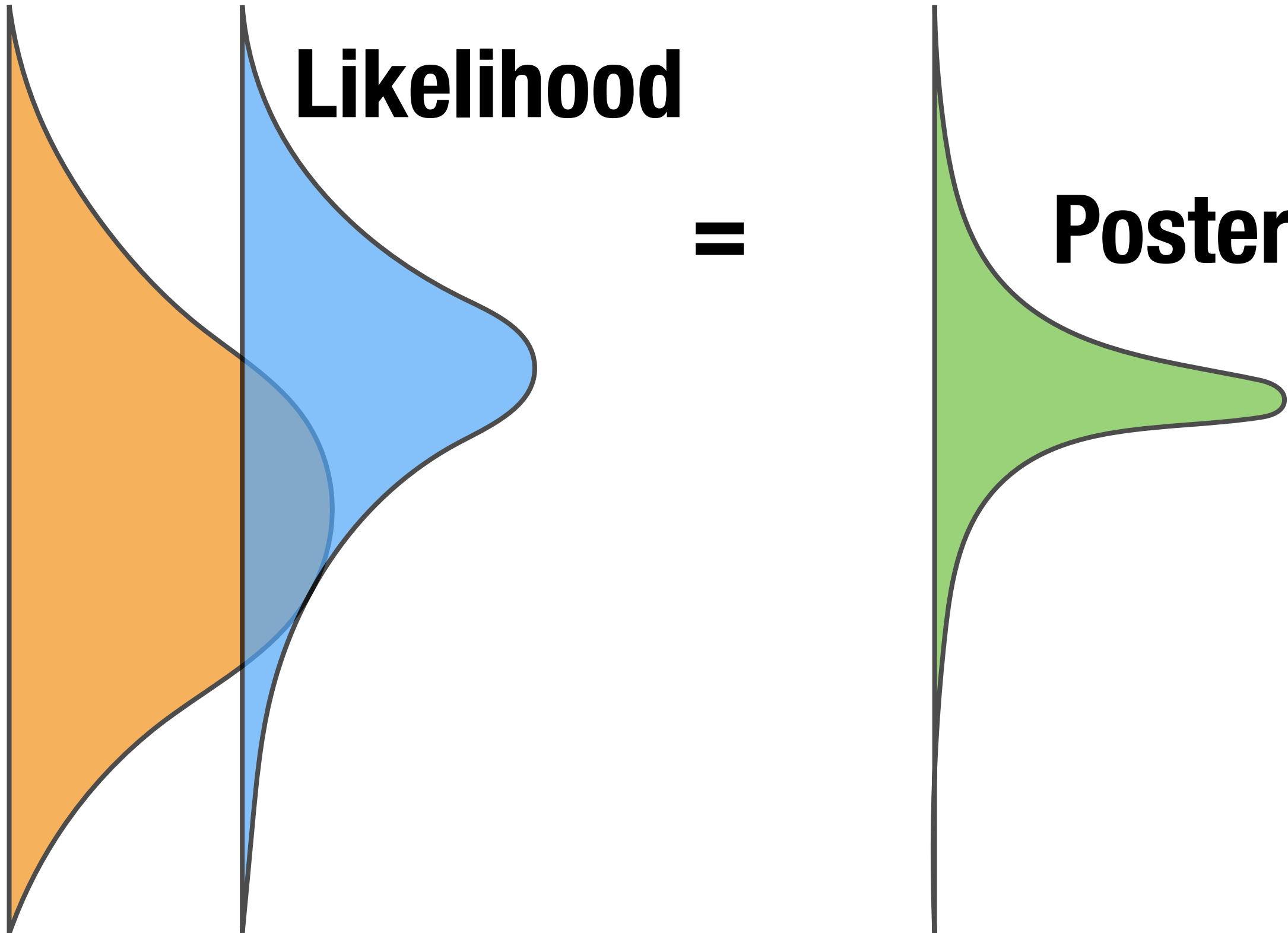
Prior

x

Likelihood

=

Posterior



BAYSPAR and BAYSPLINE

(or any type of Bayesian regression designed for paleoclimate!)
involves two applications of Bayes' Rule.

Step 1: estimate parameters: beta coefficients and tau^2 error term:

$$p(\beta, \tau^2 | \text{proxy}, \text{SST}) \propto p(\text{proxy} | \text{SST}, \beta, \tau^2) \cdot p(\beta, \tau^2).$$

Step 2: estimate SST from new proxy data. Need to invert Step 1:

$$p(\text{SST}_{new} | \text{proxy}_{new}) \propto p(\text{proxy}_{new} | \text{SST}, \beta, \tau^2) \cdot p(\text{SST}_{new}).$$

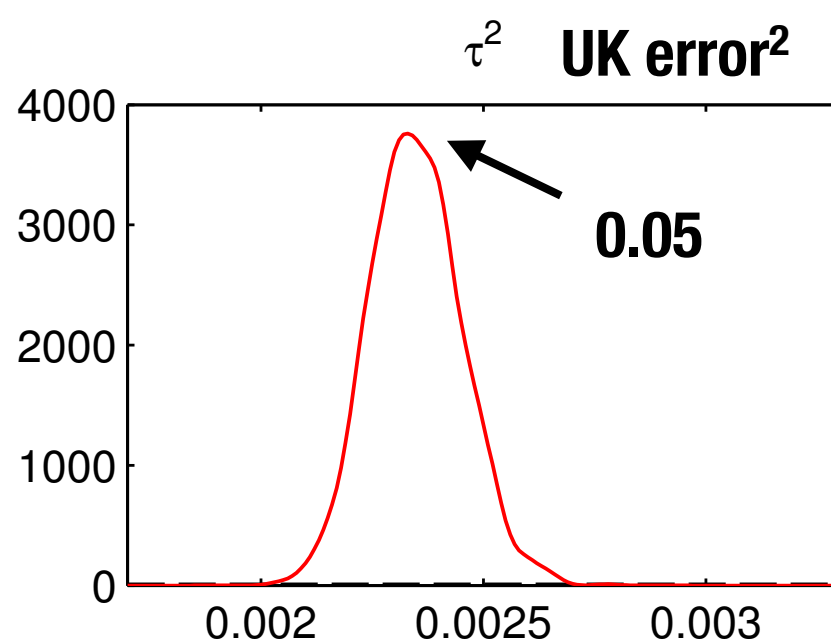
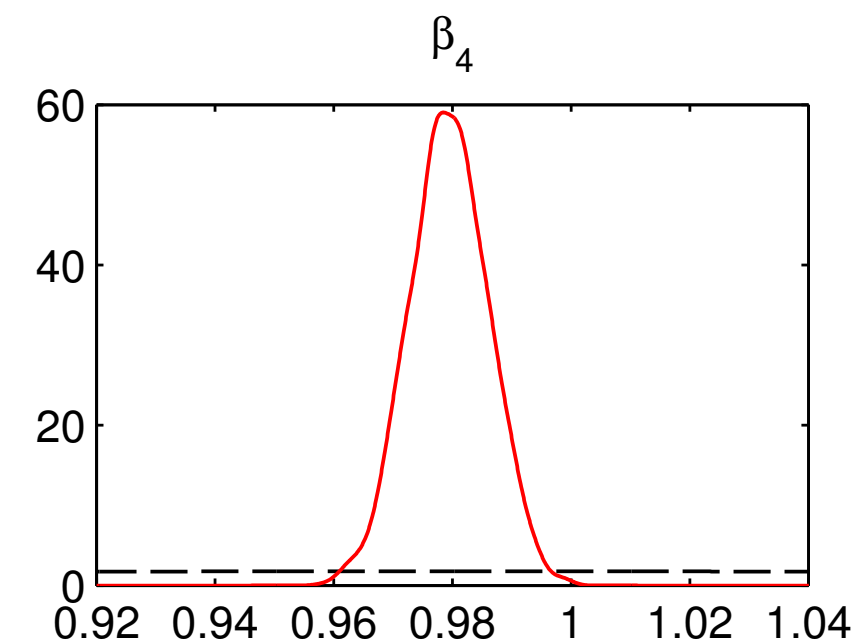
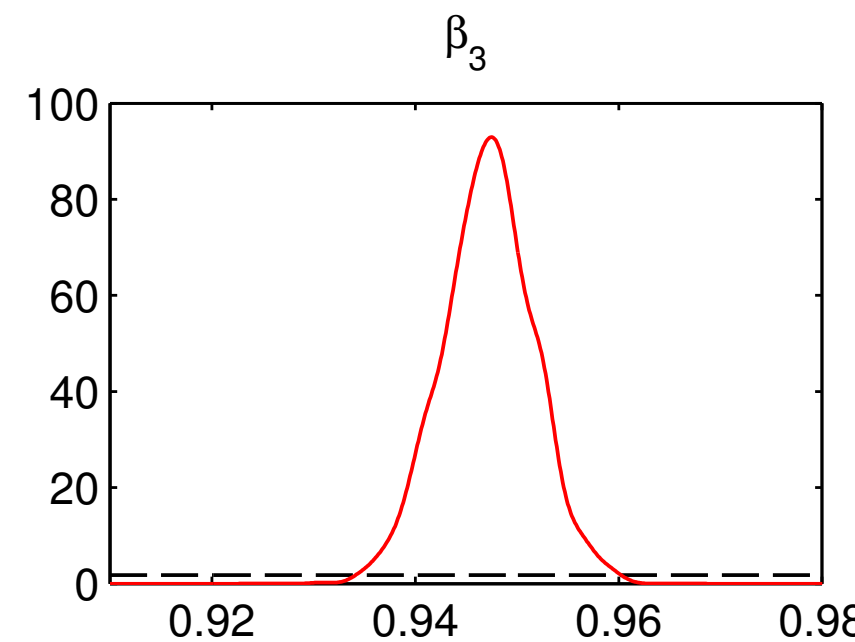
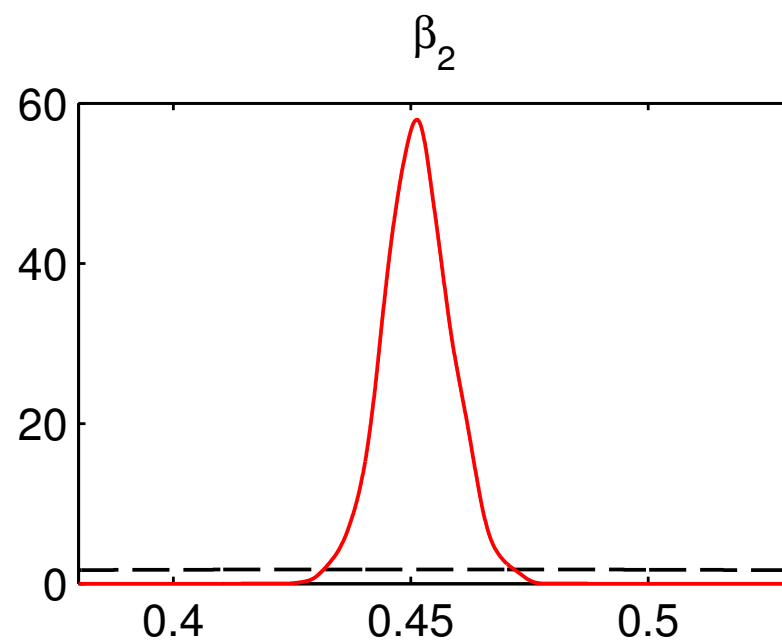
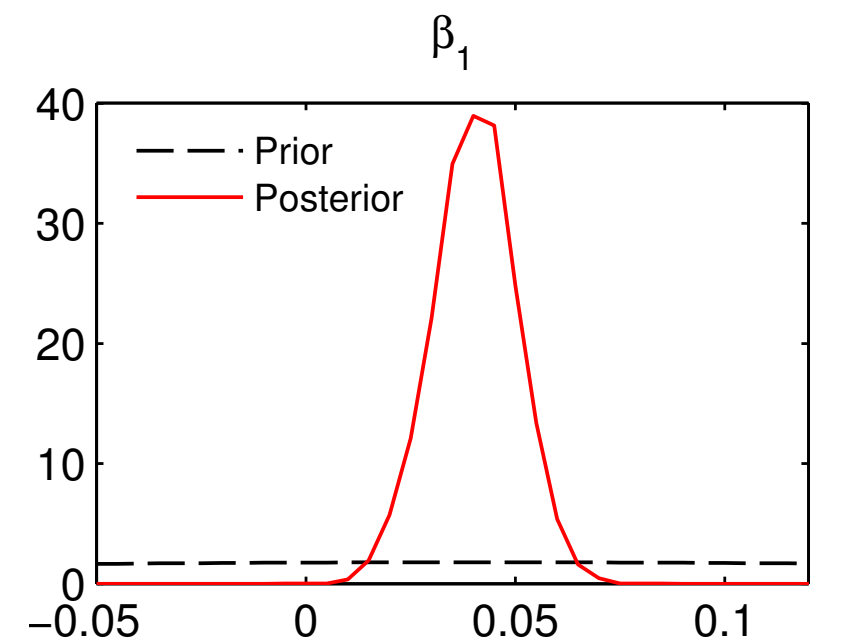
Matlab code:

bayes_regression_demo.m

predictfromreg_demo.m

Step 1: Parameter estimation

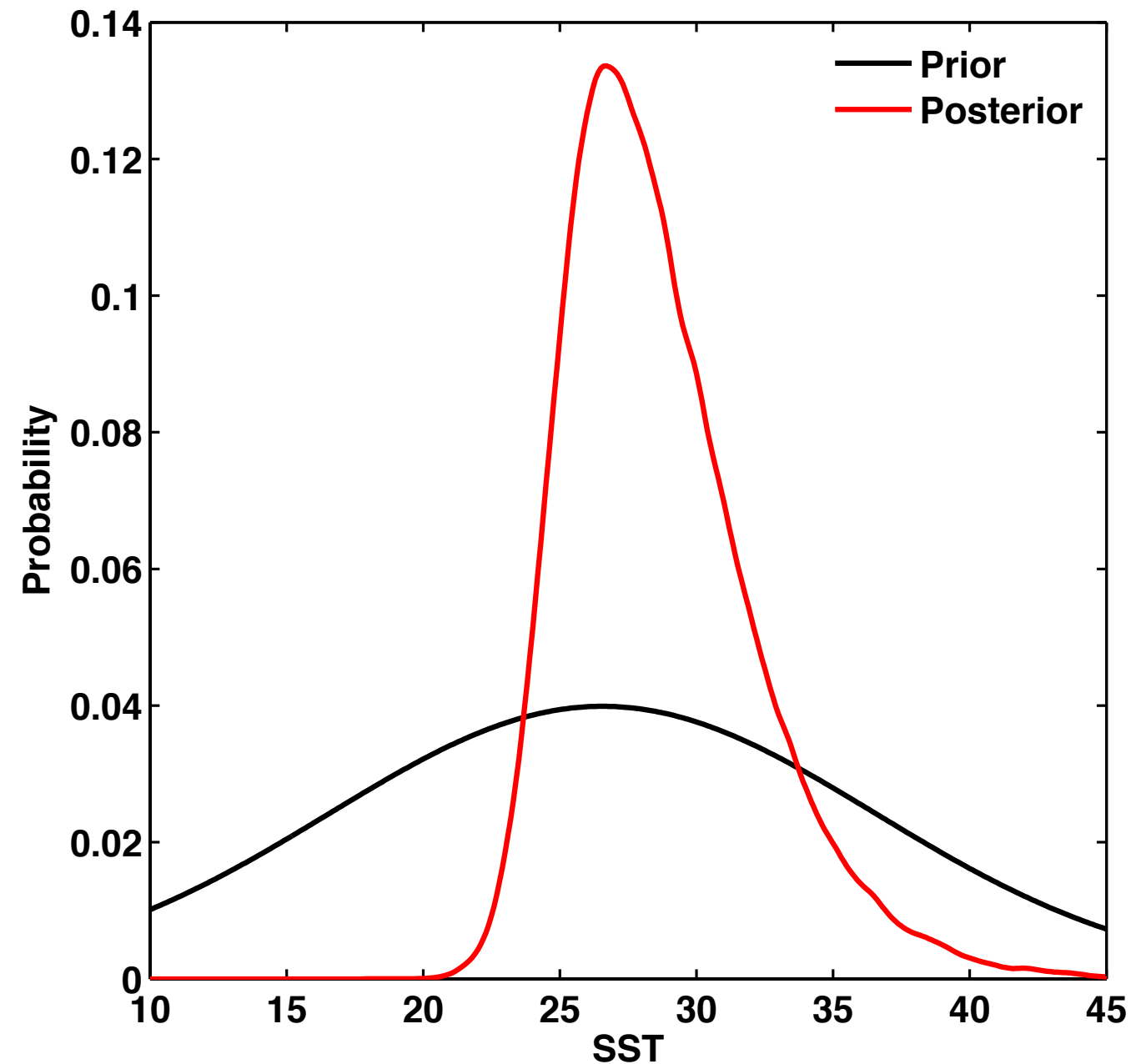
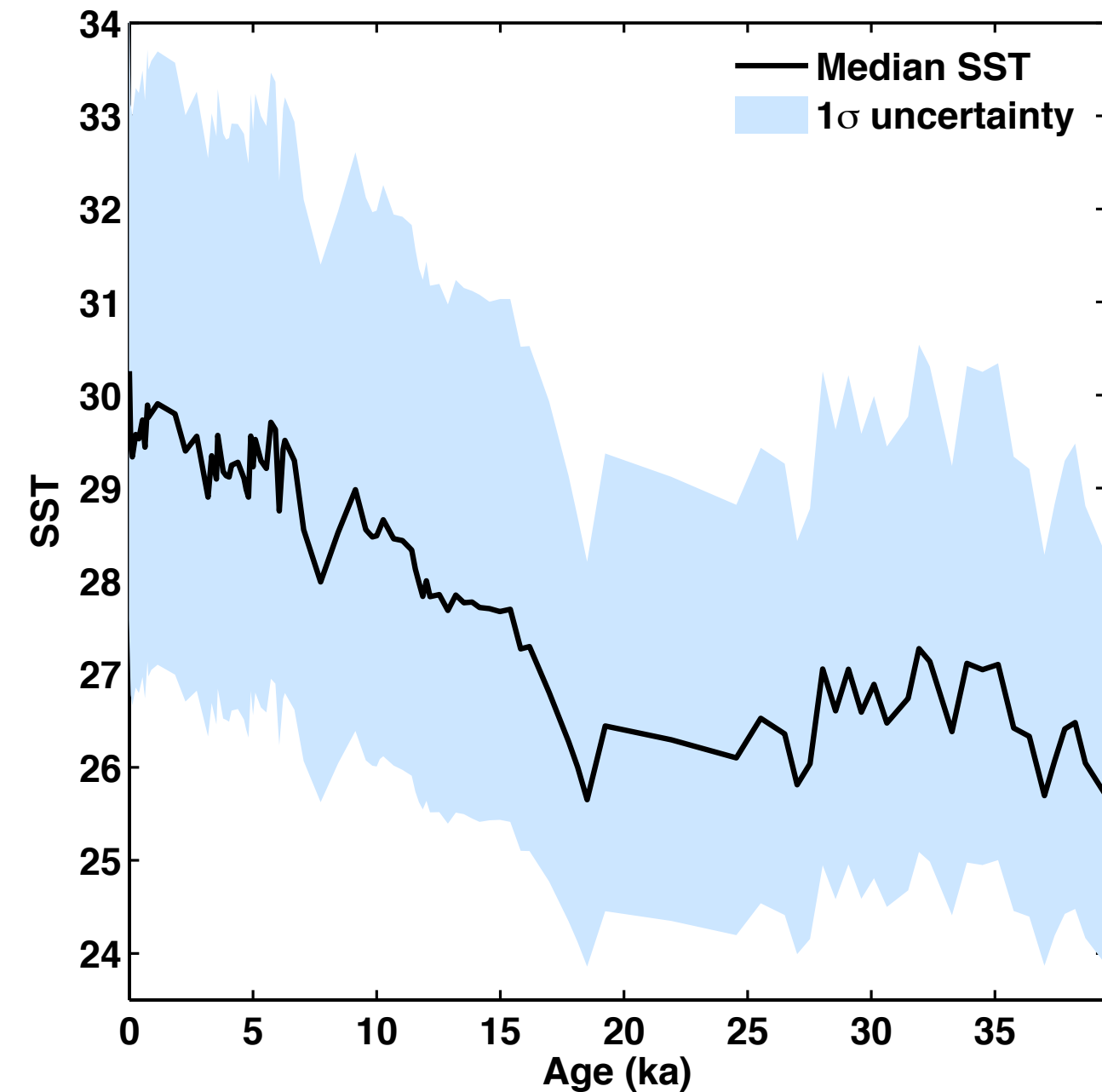
Example from BAYSPLINE. Priors are really flat,
which means likelihood dominates.



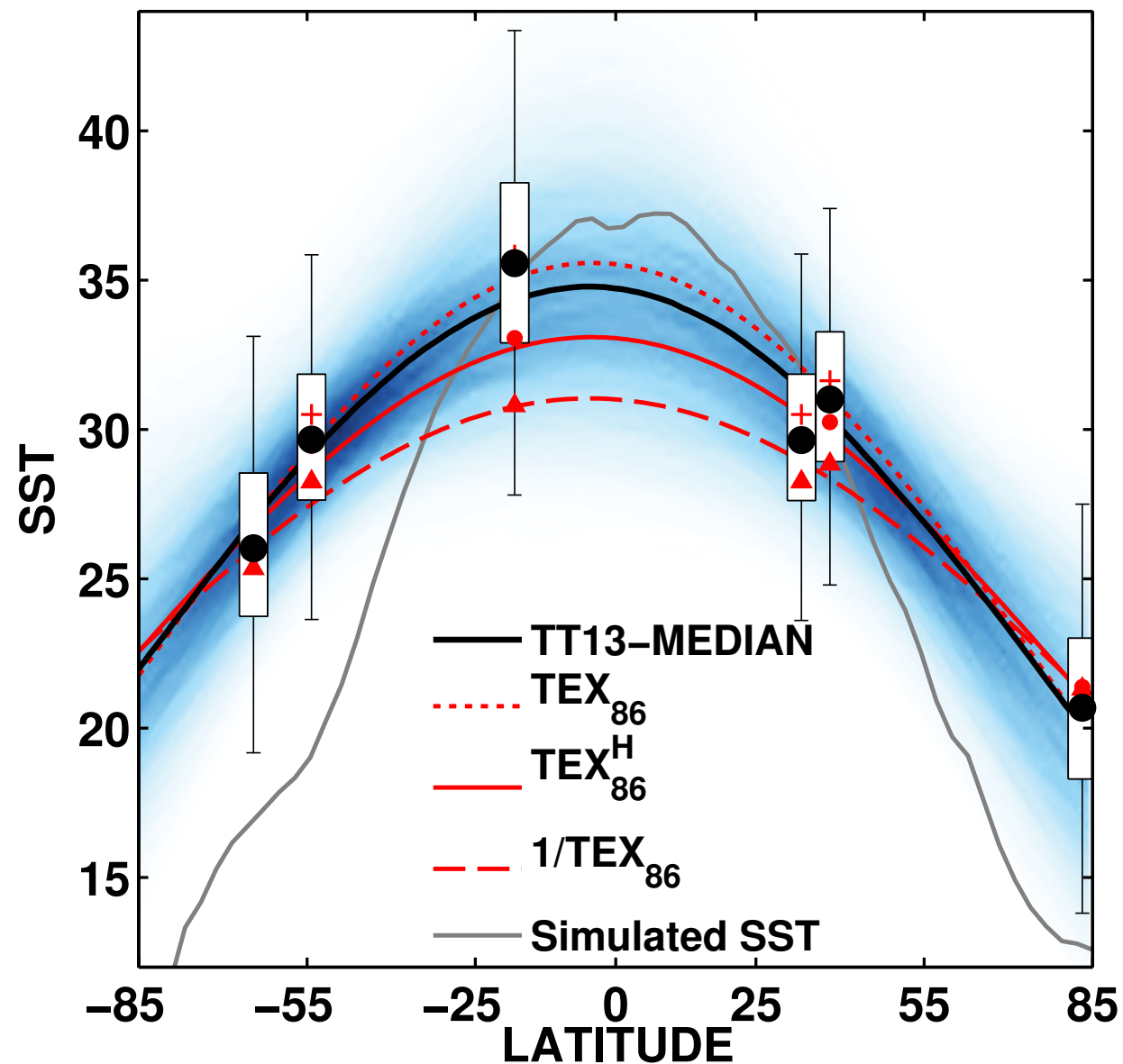
$$U_{37}^{K'} = X(SST) \cdot \beta + \varepsilon$$
$$\varepsilon = N(0, \tau^2)$$

Step 2: Temperature estimation

Example: applying BAYSPLINE to UK37 data from the Gulf of Aden.



Probabilistic Statements



There is a high (78%) probability that the pole-to-equator gradient exceeded 10°C, but it is unlikely (14%) that it exceeded 20°C.

There is a very high (93%) probability that tropical SSTs exceeded 30°C, but it is unlikely (6%) that they exceeded 40°C.

Envisioning a hierarchical model for sedimentary records

Level 1: Spatiotemporal process model
(describes SST field)

$$SST_{t+1} - \mu = \alpha \cdot (SST_t - \mu) + \epsilon_t$$

$$\epsilon_t \sim \mathcal{N}(0, \Sigma)$$

$$\Sigma_{i,j} = \sigma^2 \exp(-\phi |x_i - x_j|)$$

Level 2: Proxy forward model, can be updated to account for age model uncertainty (Werner & Tingley, 2015, Clim. Past)

$$U_{37}^{K'} = X(SST) \cdot \beta + \varepsilon$$

$$\varepsilon = N(0, \tau^2)$$

Level 3: Archive model (bioturbation or other sedimentary features)

$$U_{37obs}^{K'} | U_{37}^{K'} = \sum_{t_n} U_{37t_1+t_n}^{K'} \cdot g(t_n) + \epsilon(t_1),$$

$$\epsilon(t_1) \sim \mathcal{N}(0, \tau^2) \text{ IID.}$$

COMING SOON

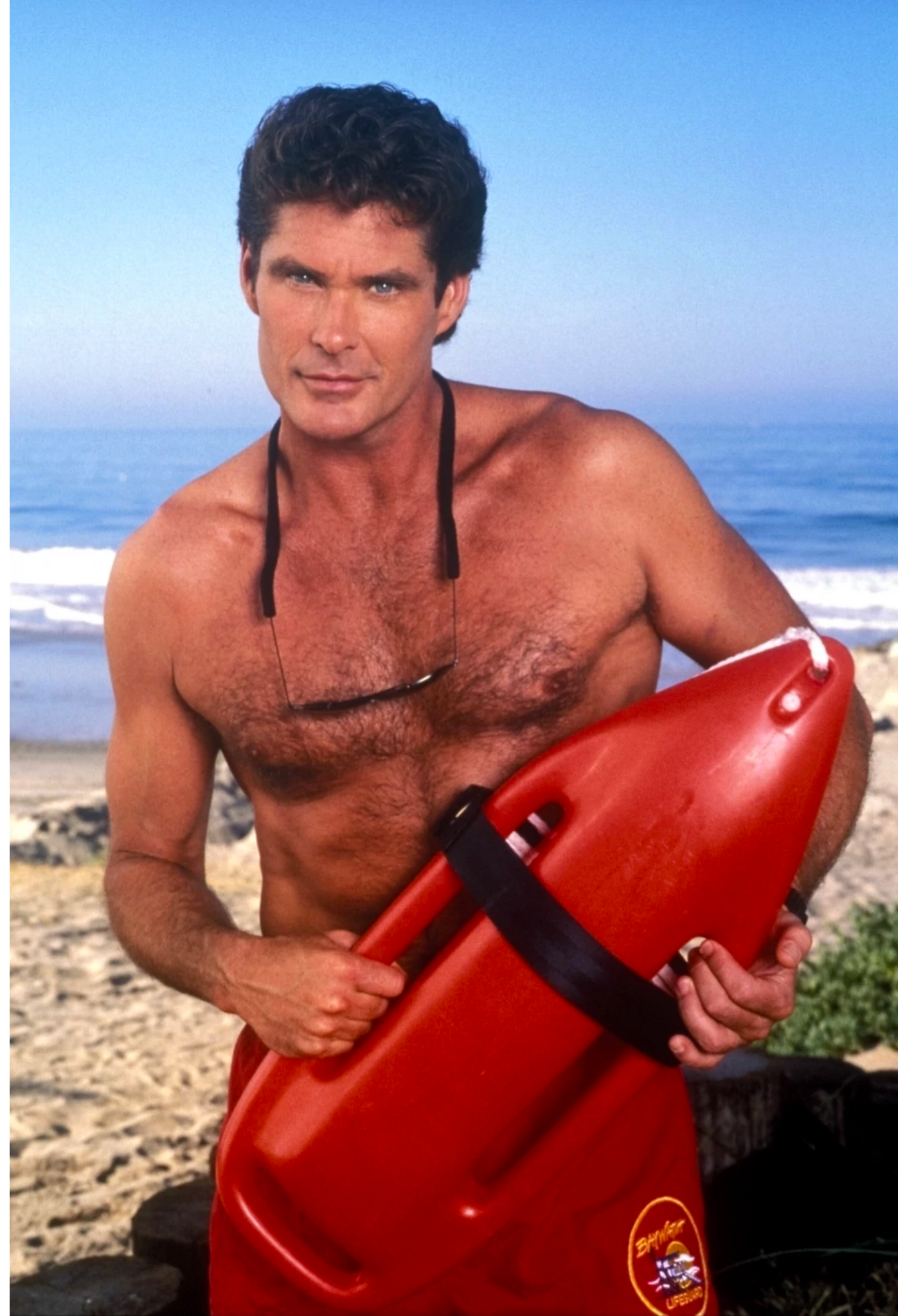
The BaySST package

BAYSPAR

BAYSPLINE

BAYFOX

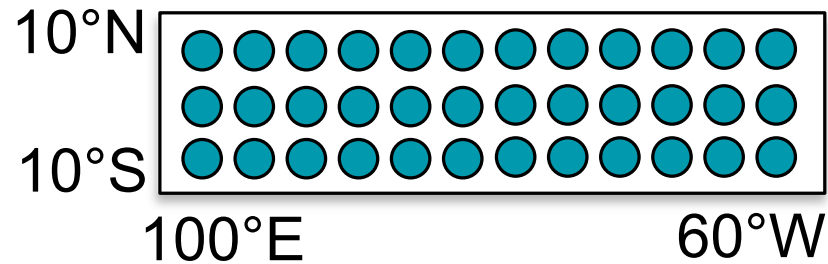
BAYMAG



Reduced space approach

After Gill et al., 2016, Paleoceanography

ERSST v5, 1854-2017 Pacific full field



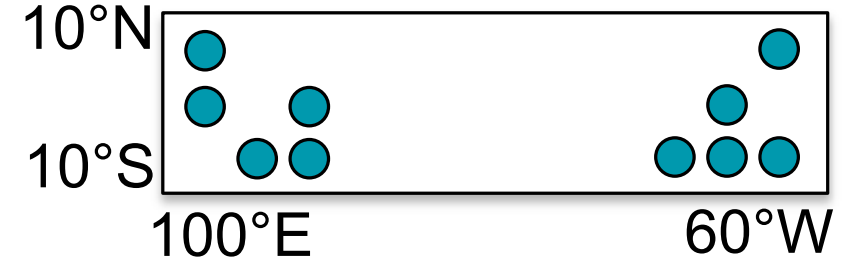
$T = N \text{ (years)} \times G \text{ (locations)}$



$$T_{[N \times G]} = U_{[N \times N]} \Sigma_{[N \times G]} V_{[G \times G]}$$

$$PCs_{[N \times G]} = T_{[N \times G]} V_{[G \times G]} \longleftrightarrow PCs_{[N \times P]} = T_{[N \times P]} V_{[P \times P]}$$

ERSST v5, 1854-2017 limited field



$L = N \text{ (years)} \times P \text{ (locations)}$



$$L_{[N \times P]} = U_{[N \times N]} \Sigma_{[N \times P]} V_{[P \times P]}$$

SVD

Regression to predict full PCs from limited PCs

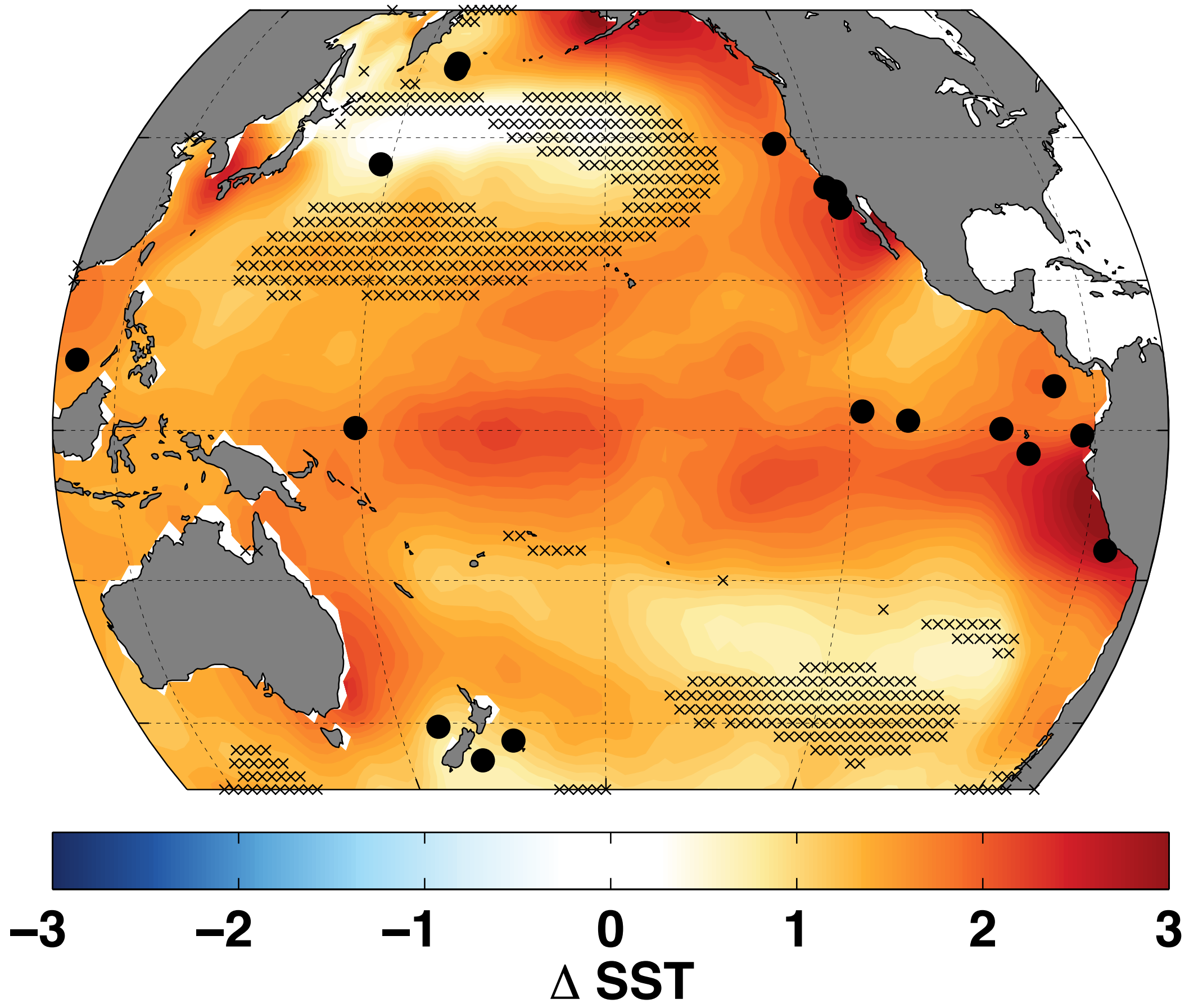
Calculate Pliocene limited field PCs: $PCs_{Plio[1 \times P]} = T_{[1 \times P]} V_{[P \times P]}$

Use regression to predict first few G PCs, rest of G is zeroes or means:

$$PCs_{Plio[1 \times G]}$$

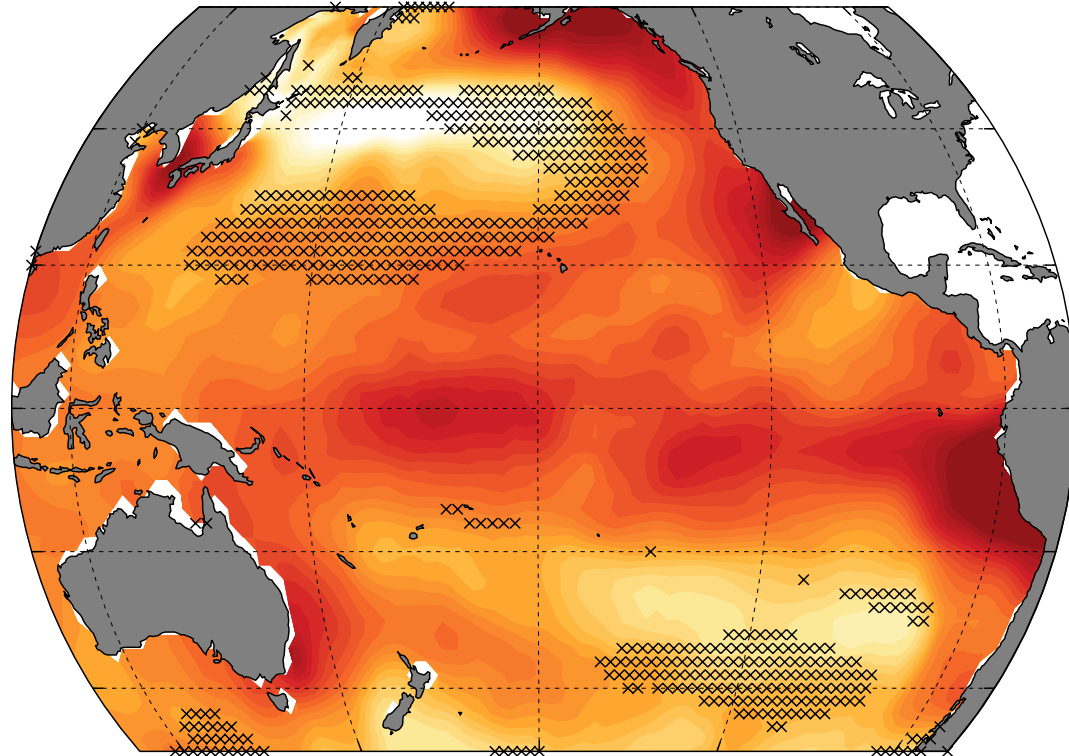
Expand back to full field: $T_{Plio[1 \times G]} = PCs_{Plio[1 \times G]} V_{[G \times G]}^T$

mid-Pliocene full field reconstruction

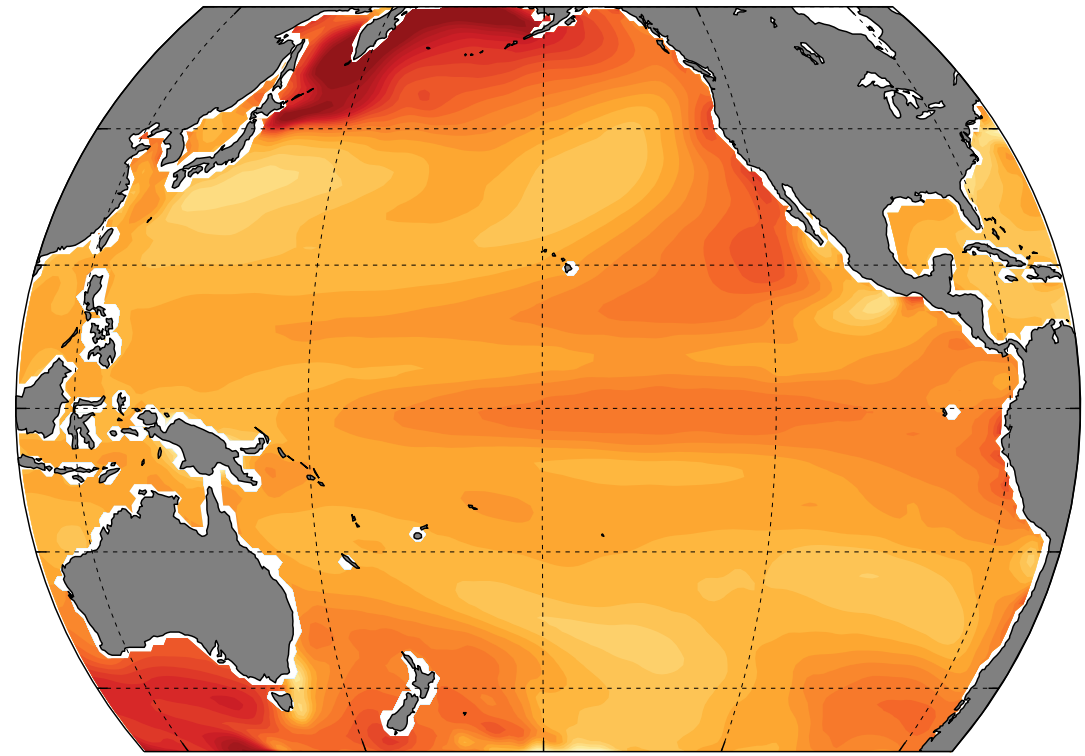


Comparison to model simulations

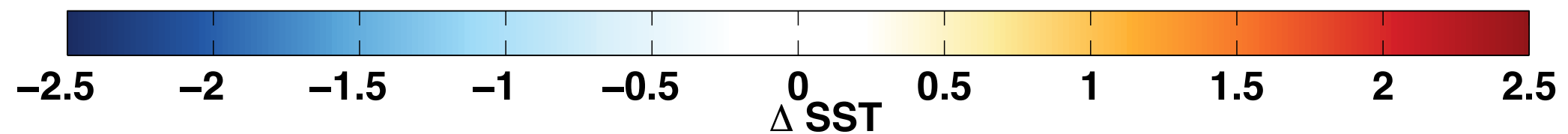
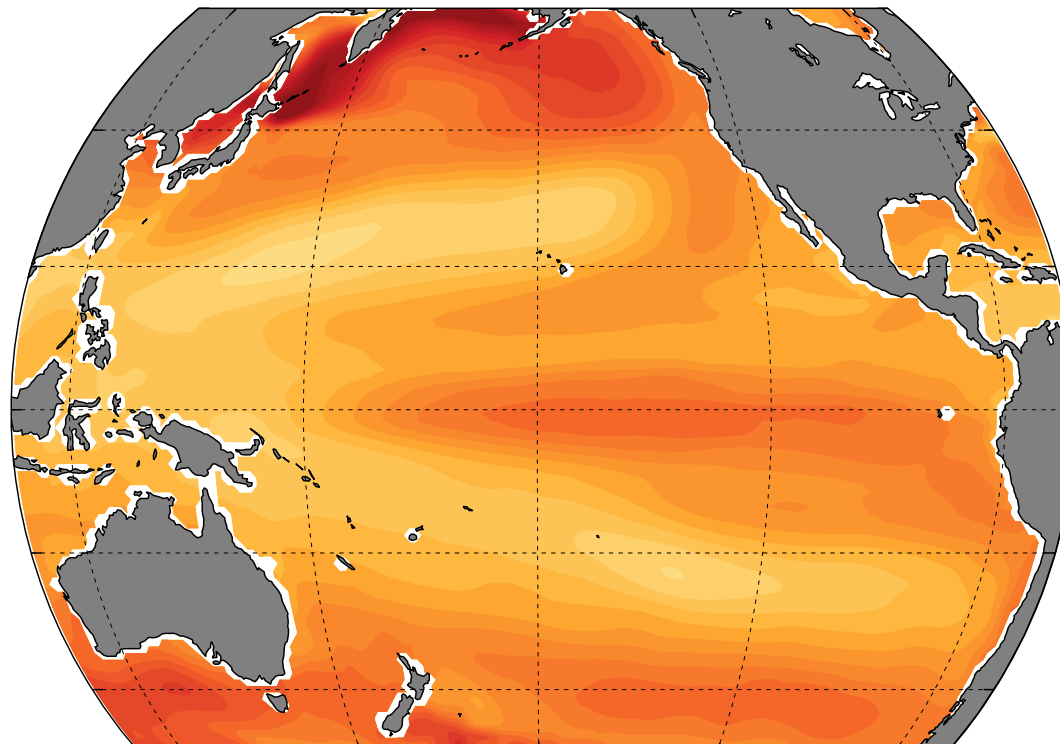
a. Pliocene Proxy Reconstruction



b. CESM1.2 Pliocene



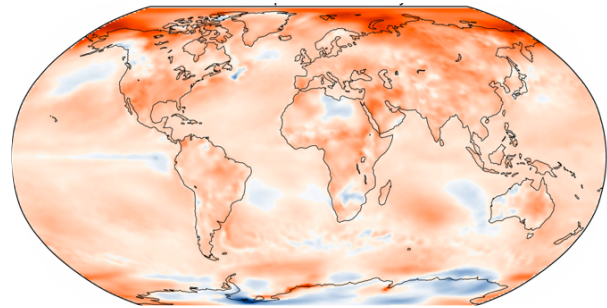
c. CESM1.2 RCP2.6



Paleoclimate data assimilation

- **Blends climate model output and proxy data.**
- **Relies on the model for the relationship between climate variables**
- **Can recover full field of many different variables**
- **Uncertainties in proxies and models can be built into the method**

Ensemble Kalman Filter Approach



$$x_a = x_b + K[y - \hat{y}]$$

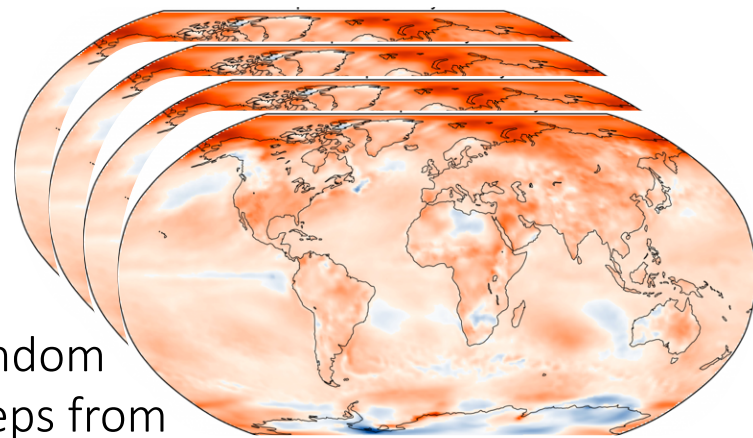
**Reconstruction
(posterior)**

Model prior

**Kalman
gain**

**Proxy
observations**

**Model estimate
of proxy
observations**



200 random
timesteps from
the 1000 year model run

$U_{37}^{K'}, Mg/Ca, \delta^{18}O$

$$U_{37}^{K'} = X(SST) \cdot \beta + \varepsilon$$

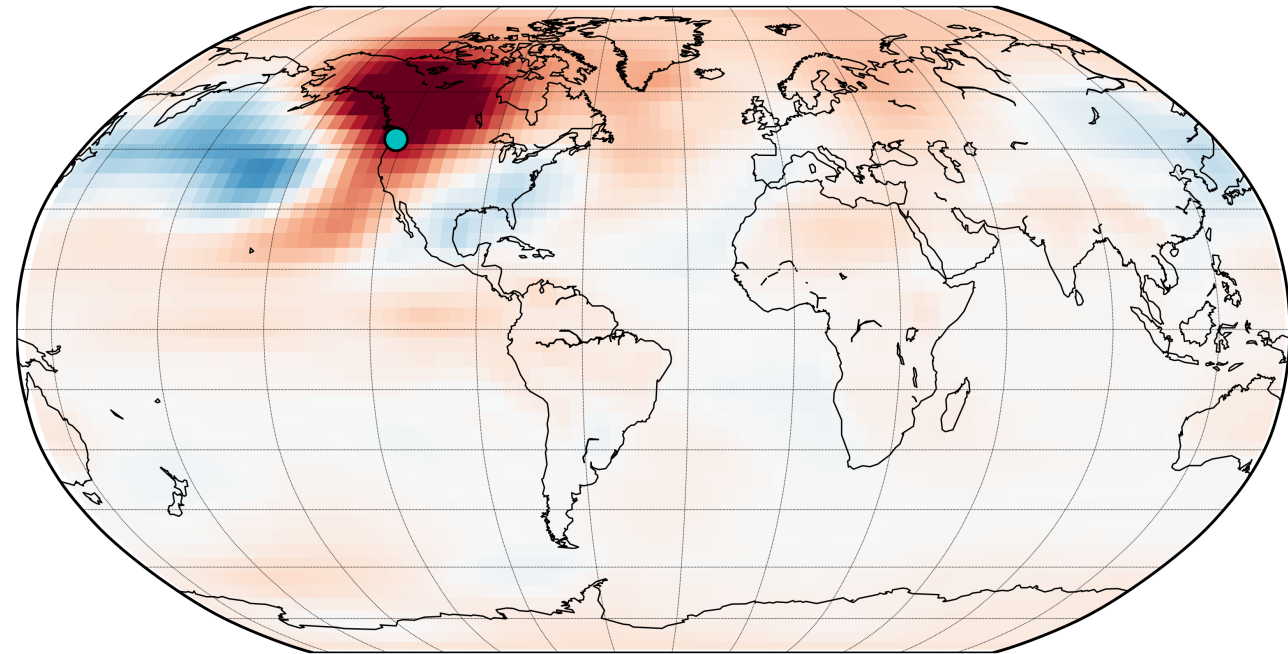
Kalman Gain

Covariance b/t the
fields in the prior
ensemble and forward
proxy estimates

$$K = \frac{cov(x_b, \hat{y})}{var(\hat{y}) + R}$$

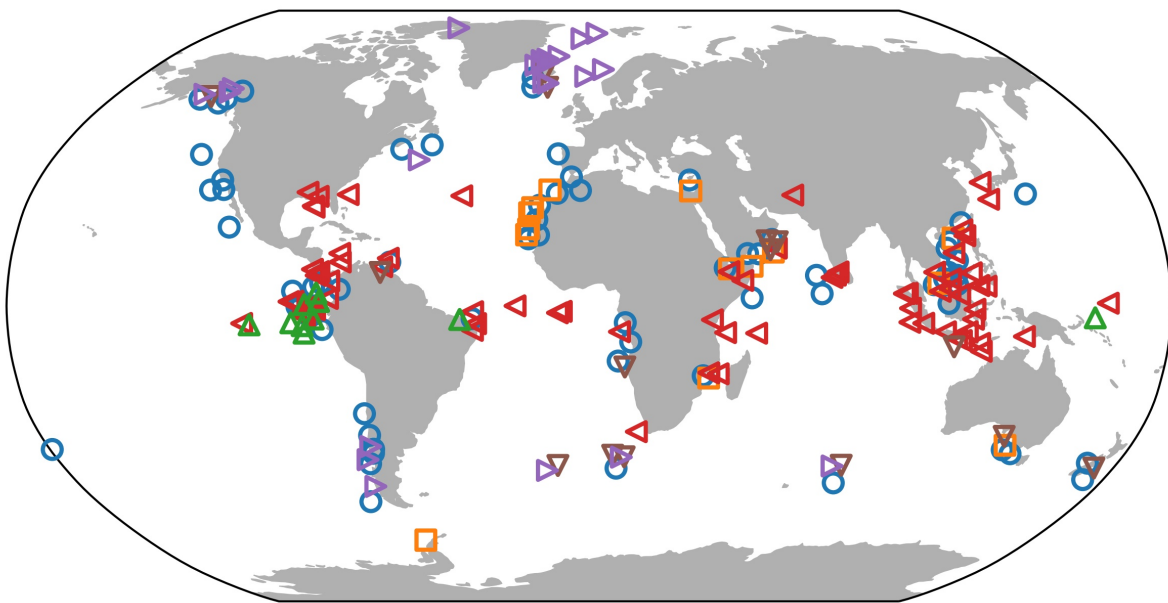
Uncertainty in the
model prior (in proxy
space)

Covariance (temperature, $\delta^{18}\text{O}$)

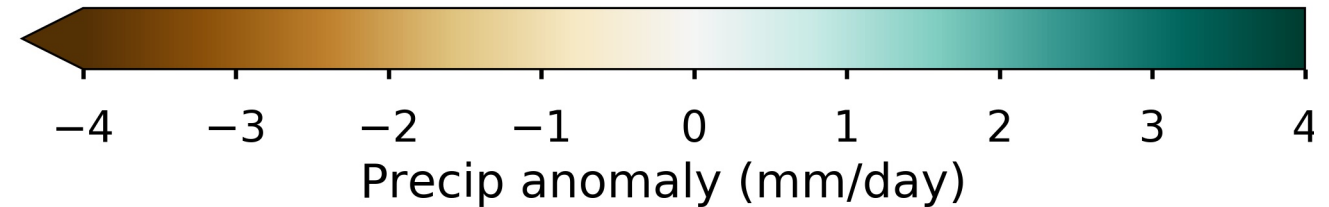
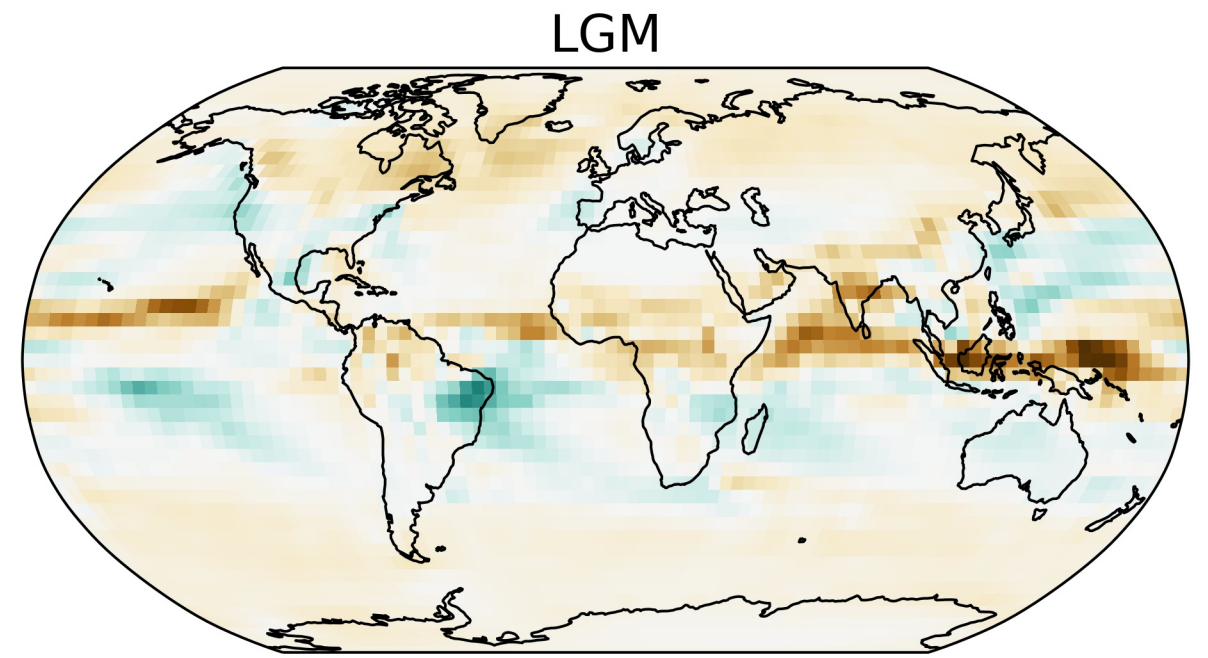
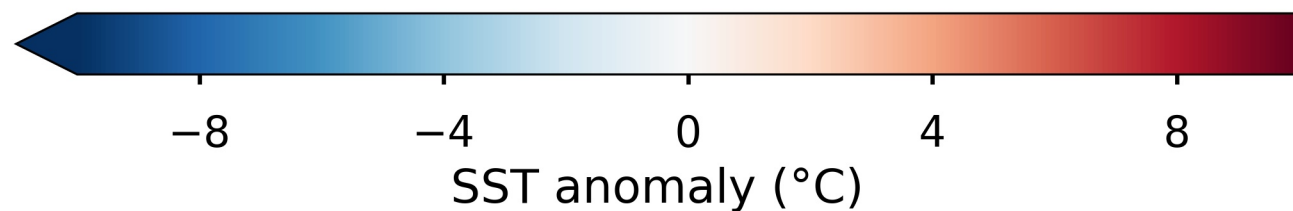
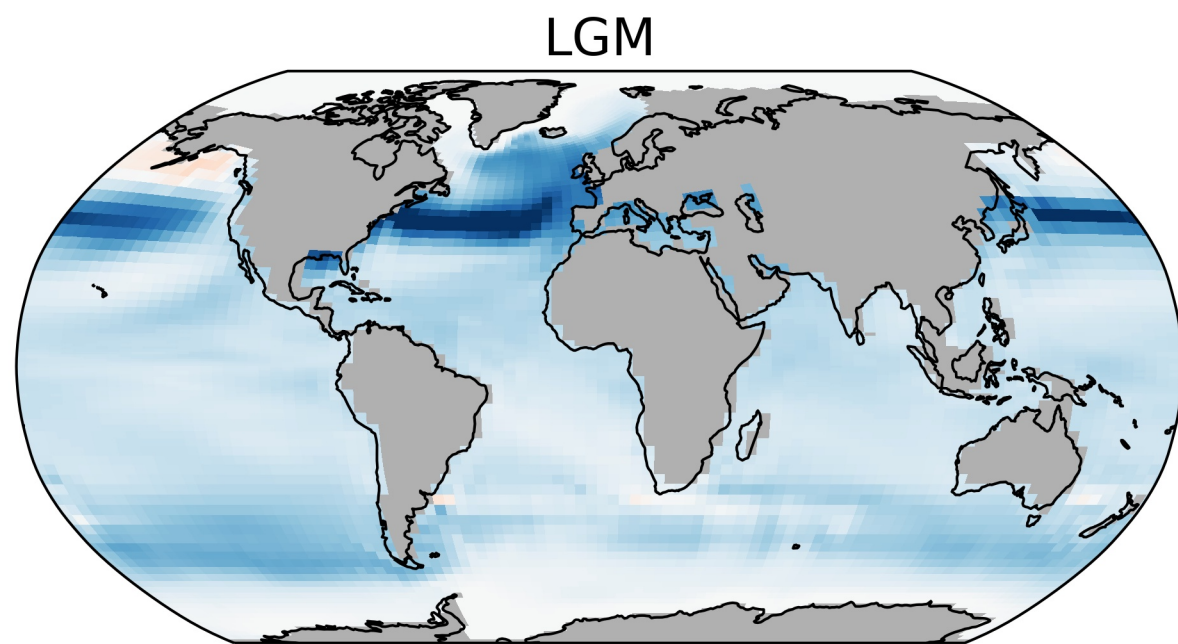


Observed proxy
uncertainty (also
a variance term)

Example: SST proxies + TraCE



- | | |
|---|--|
| ○ U_{37}^K | △ $\delta^{18}O_{G. \text{ sacculifer}}$ |
| □ TEX_{86} | ▽ $\delta^{18}O_{G. \text{ bulloides}}$ |
| △ $\delta^{18}O_{G. \text{ ruber (white)}}$ | ▽ $\delta^{18}O_{N. \text{ pachyderma (sinistral)}}$ |



Compare with model- and data-only estimates

