

Atmospheric Coupling via Energetic Particle Precipitation (EPP)

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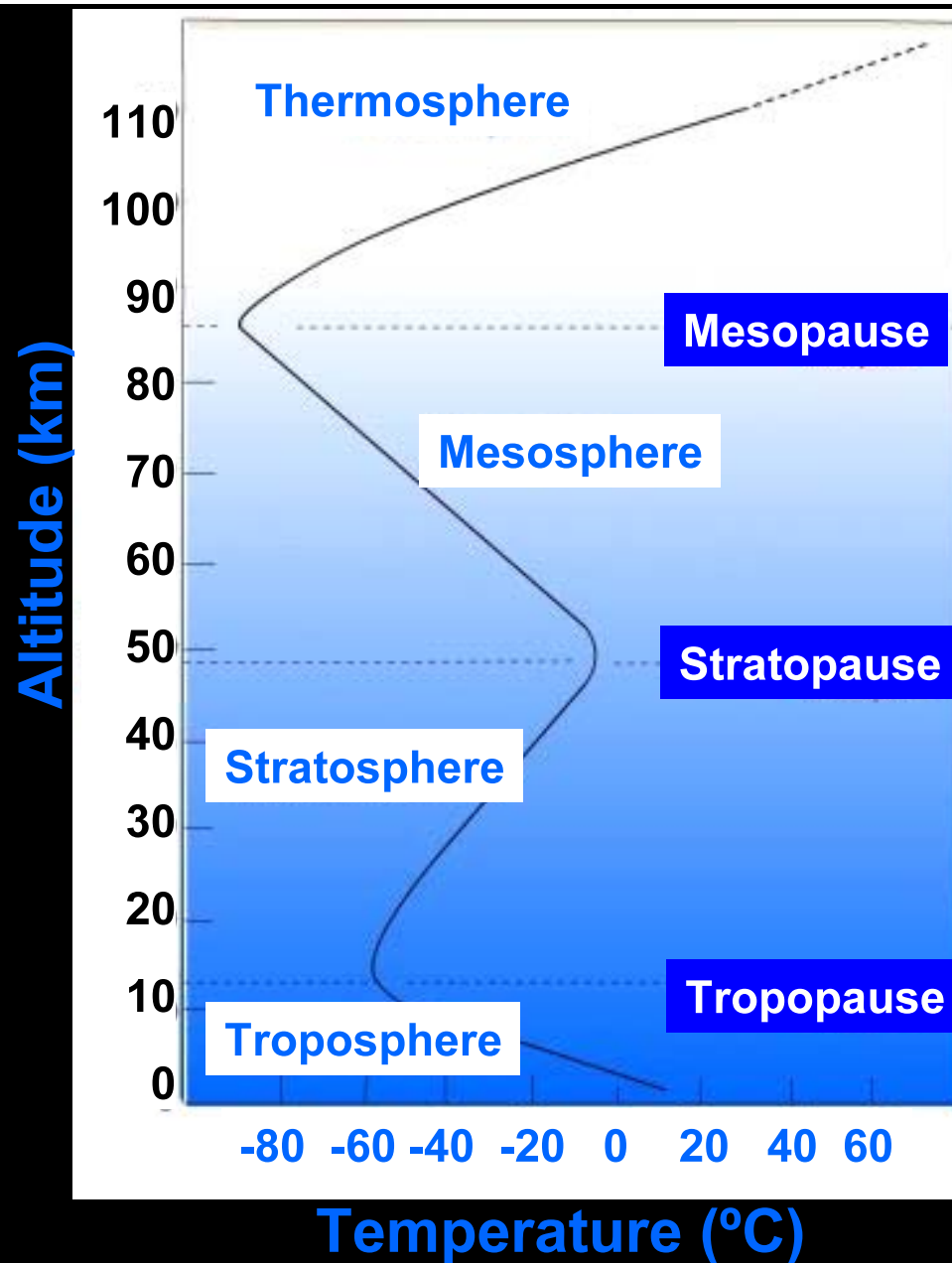
Mike Mills

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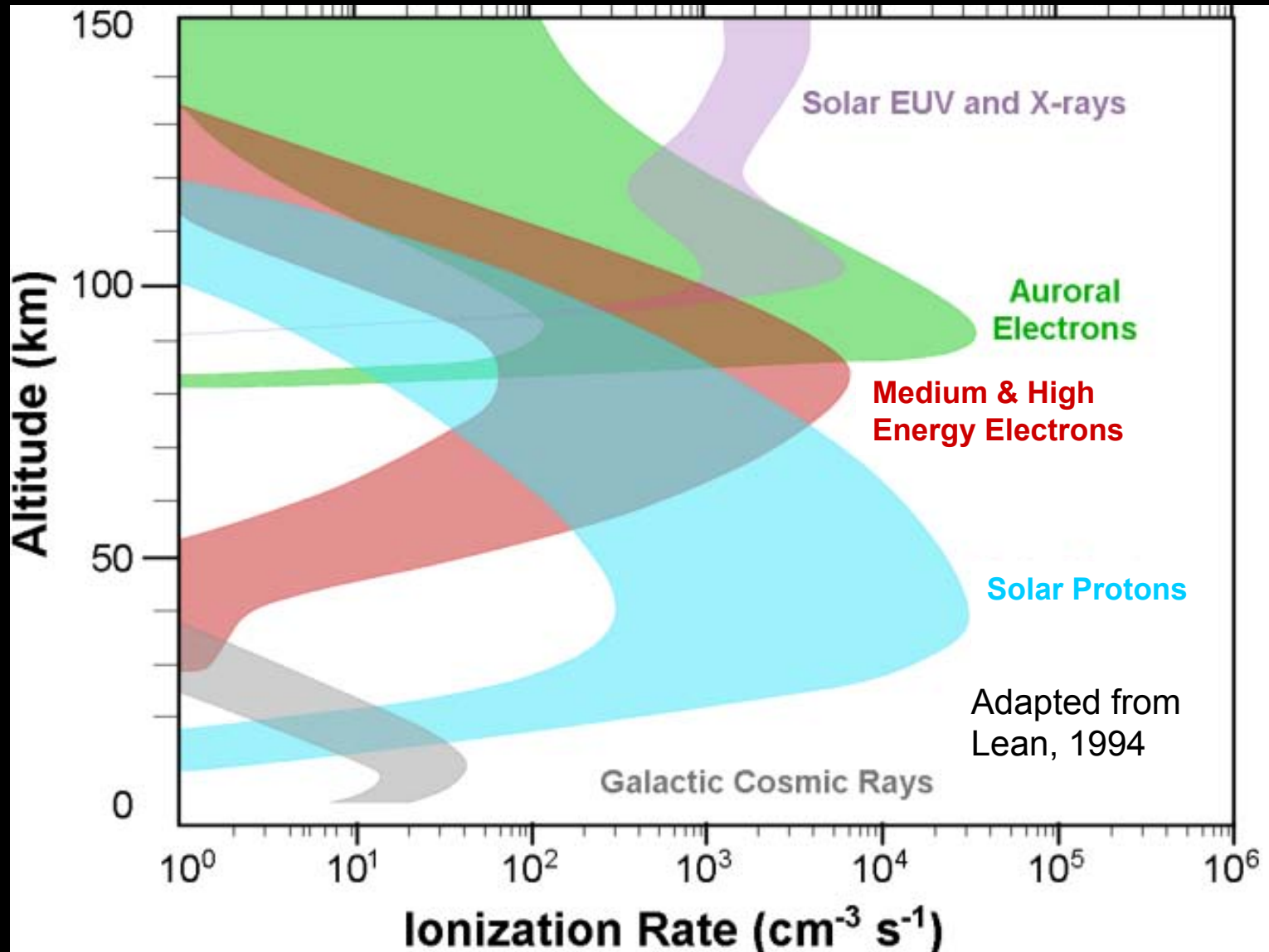
Vertical Coupling

- Troposphere to thermosphere
- In the polar region
- Mainly in winter

Outline

Intro to Mechanism
Historical Evidence
Recent Anomalies
Model Results

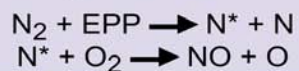
Energetic Particle Precipitation



Solar-Magnetosphere Energy

Low-energy particles
precipitate in thermosphere

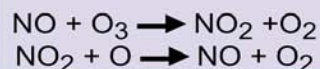
*Energetic particles
produce NO*



NO transported
downward and/or
produced locally



*NO catalytically
destroys ozone*



POLAR VORTEX

ring
current,
radiation
belts

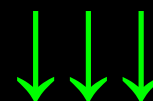
polar
cap

Strong
horizontal
winds
prevent
transport
to lower
latitudes

auroral
zone

High-energy
particles precipitate
in mesosphere and
stratosphere

Energetic Particle Precipitation (EPP)



Ionization & Dissociation



NO_x and HO_x

NO_x and HO_x Destroy Ozone

EPP Effects on the Stratosphere

DIRECT EFFECT (DE): NO_x *produced in* stratosphere.

- Requires highly energetic particles

Stratosphere: > 300 keV electrons
 > 30 MeV protons

- Sporadic production

INDIRECT EFFECT (IE): NO_x produced in mesosphere or thermosphere and *descends to* stratosphere

- Requires only medium or low energy particles
- Routine production



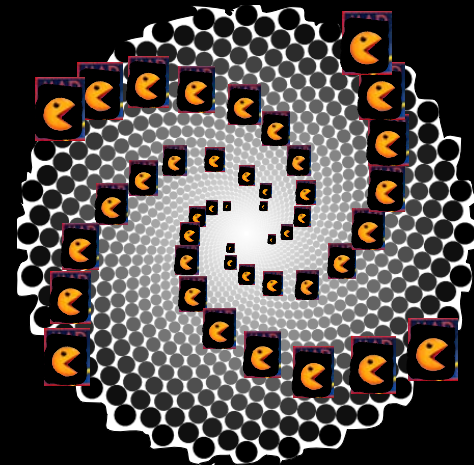
EPP Indirect Effect

NO produced in mesosphere or thermosphere and *descends to* stratosphere

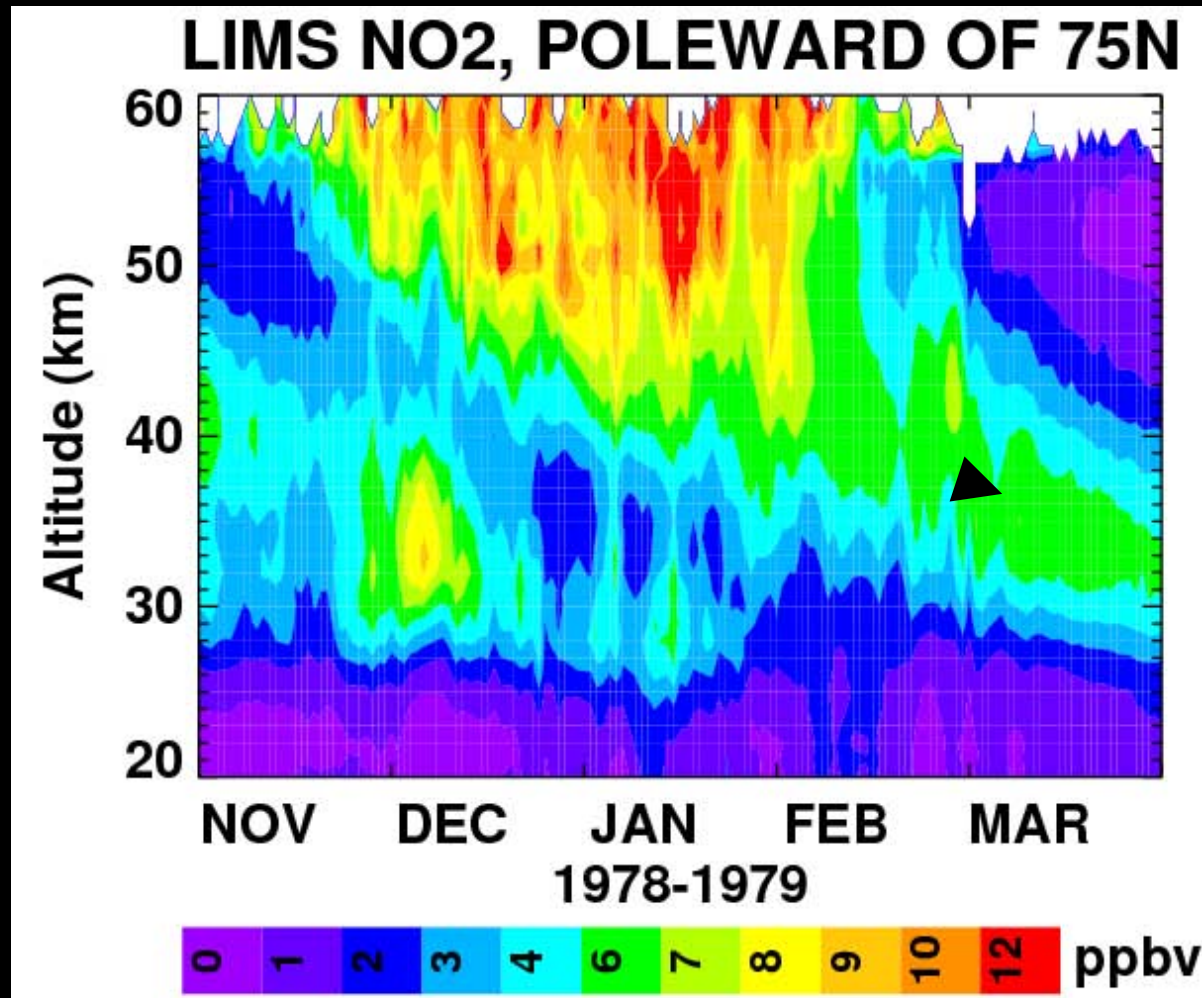
Routine production, but MLT NO_x has lifetime of days, so stratospheric effects require efficient downward transport

- Polar night
- Strong descent in MLT
- Confinement in Polar Vortex

NO_x in stratosphere destroys ozone



First satellite observations of EPP Indirect Effect from LIMS in NH, 1978-1979



**EPP is the
ONLY source of
mesospheric
NO_x in the polar
winter!**

Based on Russell
et al., 1984



Satellite observations of EPP-NO_x were sparse from 1979 to 2003.

EPP Investigations mainly used solar occultation data from SAGE, HALOE, and POAM

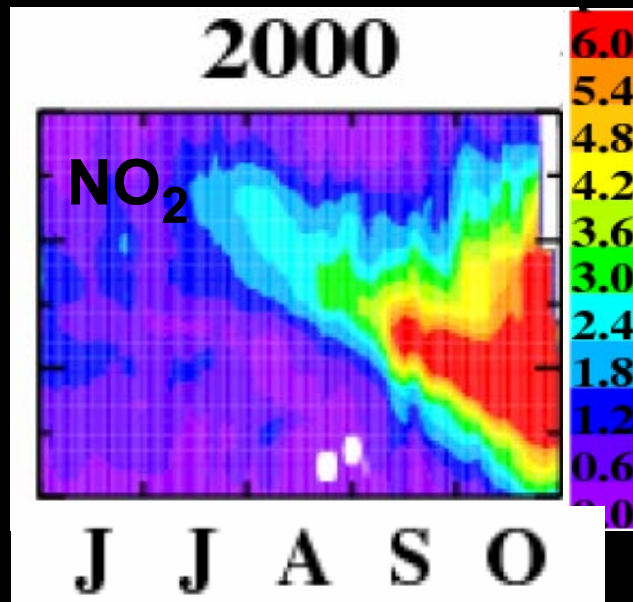
- **Sparse geographic coverage**
- **SAGE & POAM only measure NO₂, not NO**
- **No polar night**

In 2003 more data became available

**e.g., GOMOS, MIPAS, SCIAMACHY, ACE-FTS
(Still cannot measure polar night NO in MLT)**



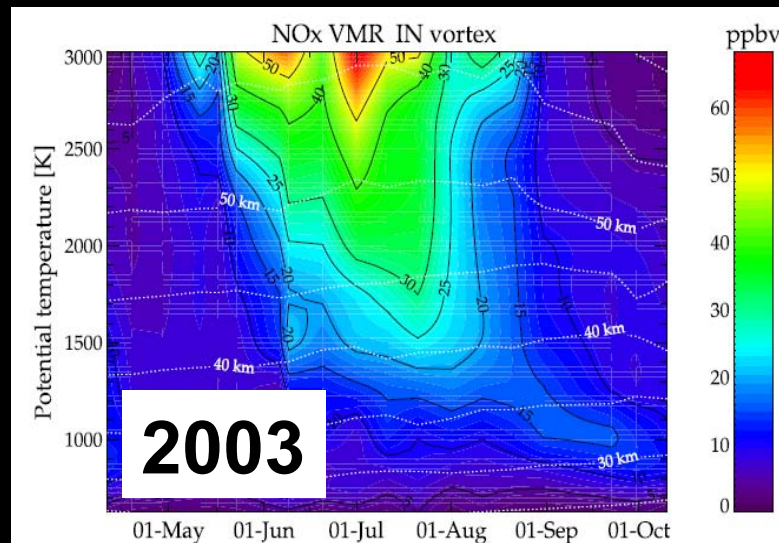
POAM: SH



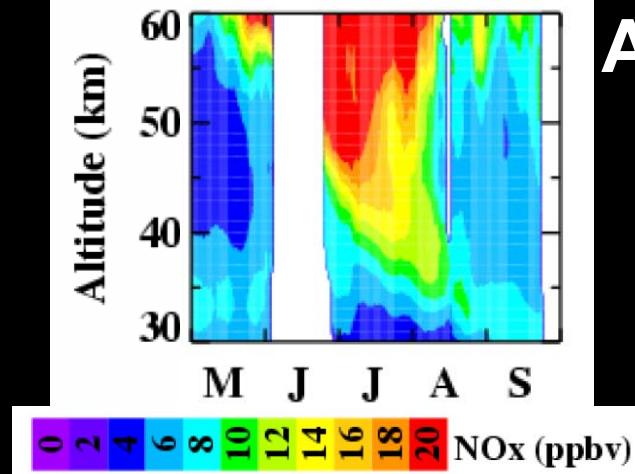
Randall et al., 2007

MIPAS: SH

Funke et al., 2005



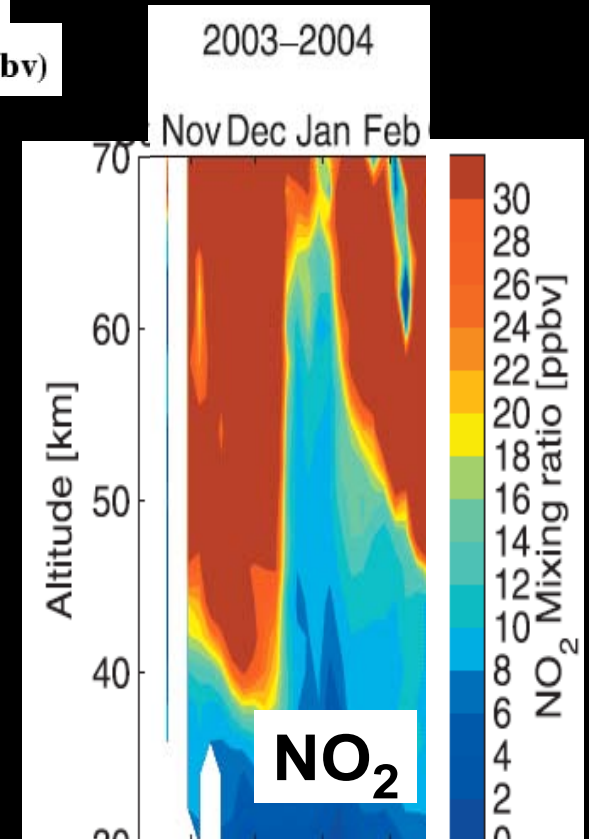
2005



Randall et al., 2007

ACE-FTS: SH

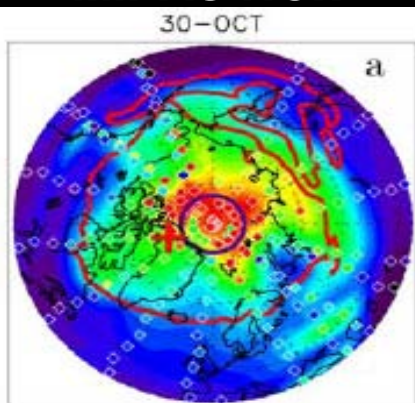
GOMOS: NH



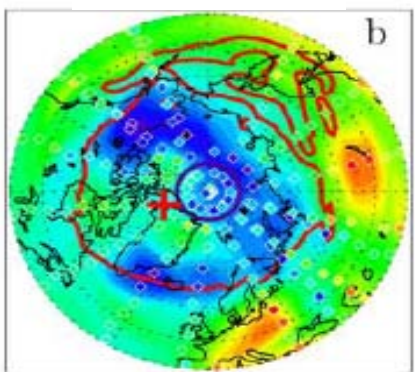
Seppala et al., GRL 2007

EPP-NO_x enhancements accompanied by ozone reductions

MIPAS NO_x

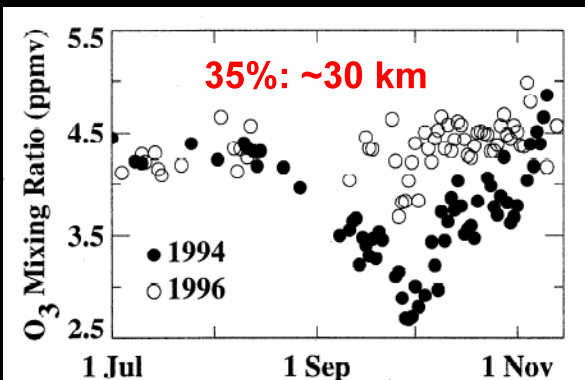


MIPAS O₃



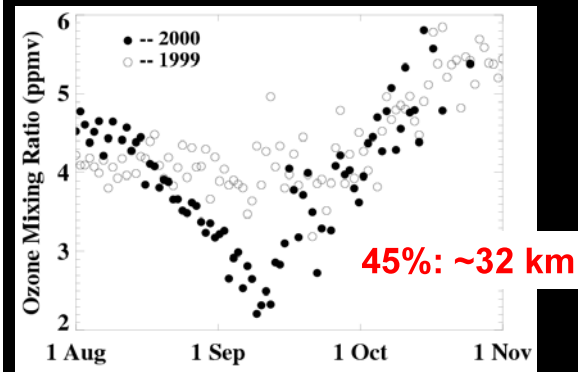
Lopez-Puertas et al., 2005

POAM ozone: 1994 vs. 1996

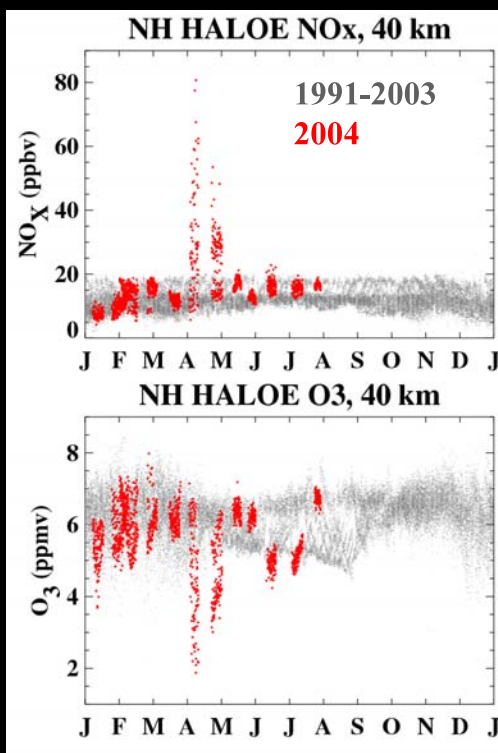


From Randall et al., 1998

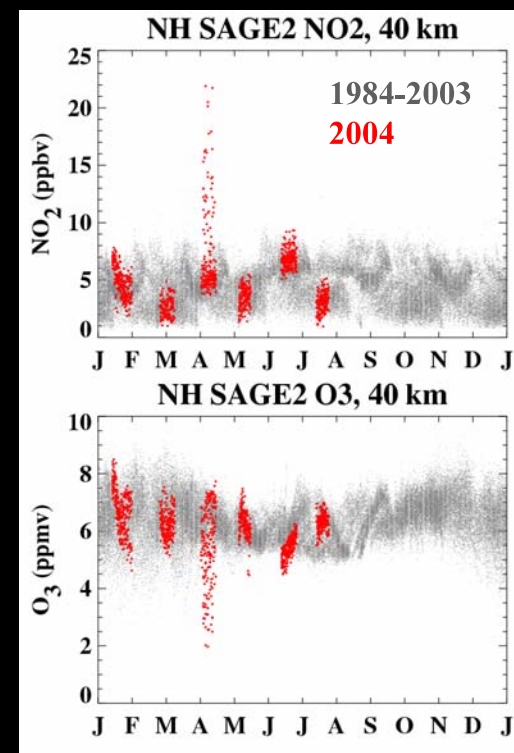
POAM ozone: 1999 vs. 2000



From Randall et al., 2001

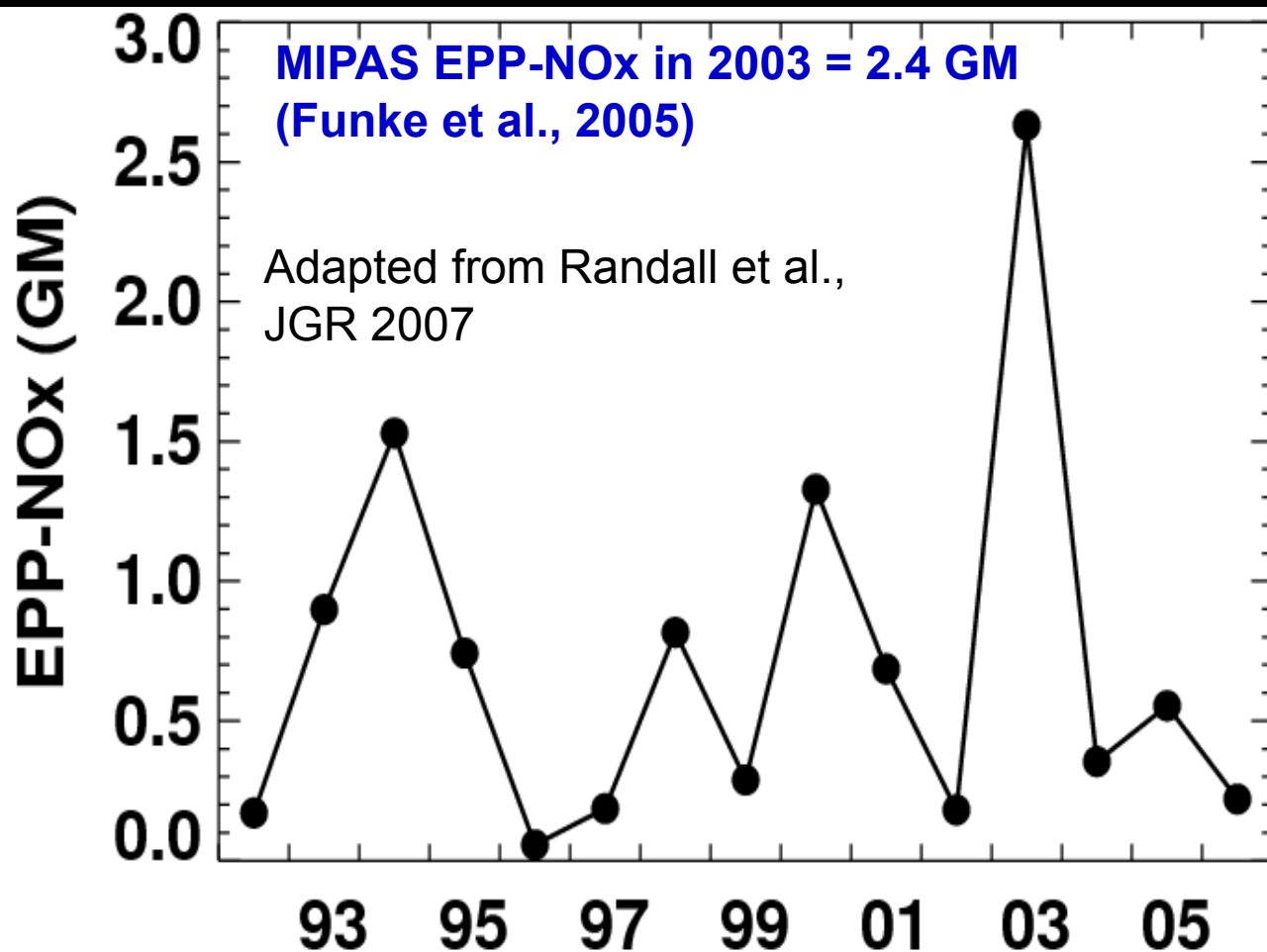


Randall et al., 2005



Randall et al., 2005

EPP-NO_x entering the *Southern Hemisphere* Stratosphere

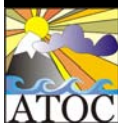


How
significant is
this?

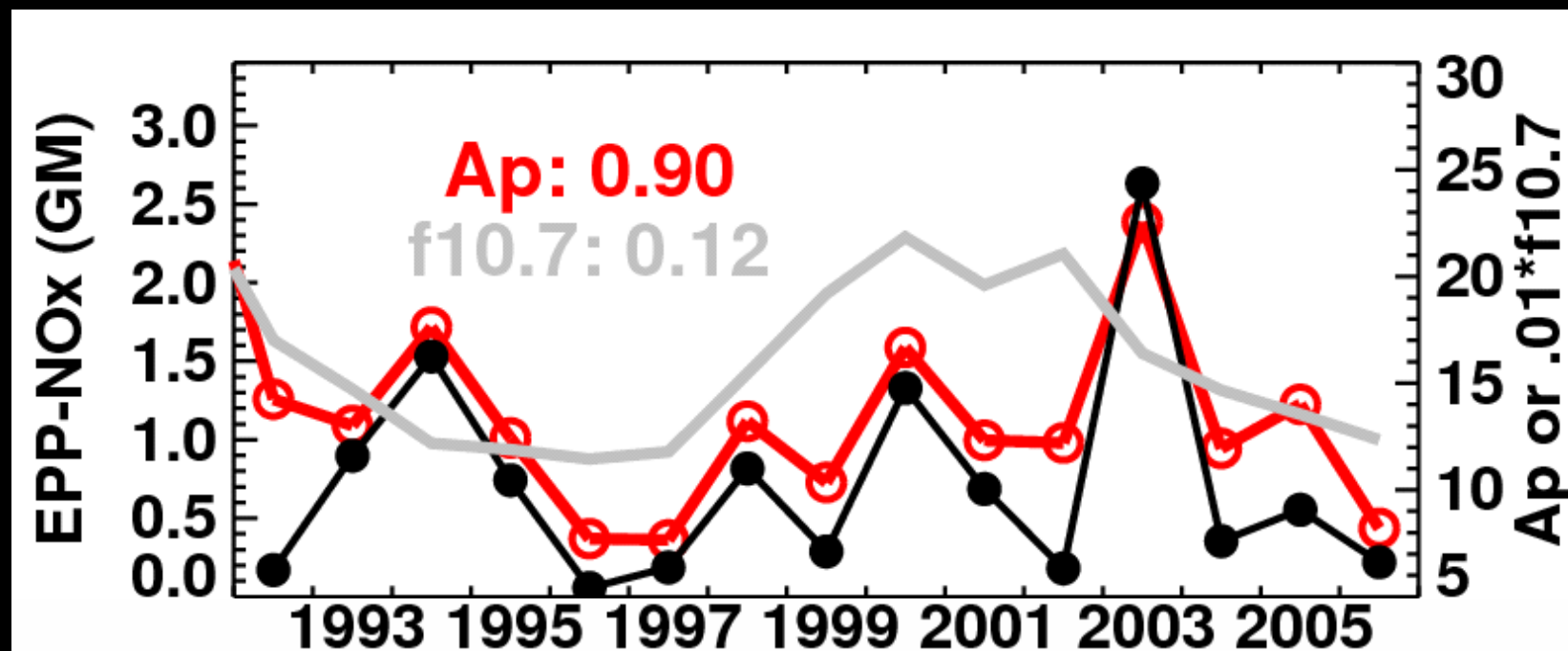
Up to 10% of
total SH NO_y
(rest from N₂O
oxidation)

Up to 40% of
polar NO_y

What determines the interannual variability?



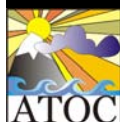
Correlation of EPP NO_x entering the SH stratosphere with Ap and Solar f10.7 (Apr-Aug)



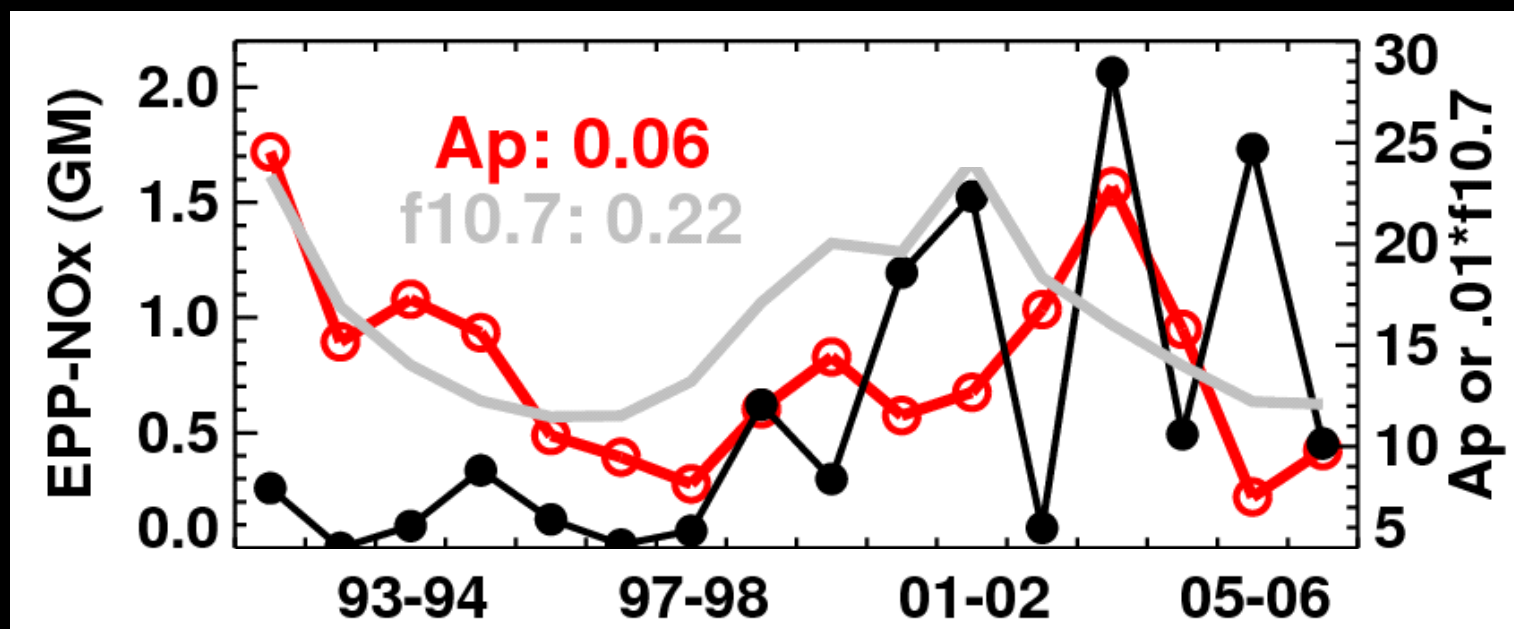
Adapted from Randall et al., JGR 2007

Strong correlation with Ap (and auroral & MEE power, and thermospheric NO) but not F10.7

Variability in SH stratospheric NO_x from EPP controlled by variation in EPP- NO_x production: Stable Dynamics!



EPP-NO_x entering the *Northern Hemisphere* Stratosphere vs. Ap & F10.7

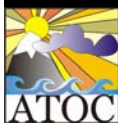


In NH, correlations with Ap index & F10.7 are poor.
Both dynamical variability and EPP play critical roles in controlling the NH variability

2001-2002: SPE

2003-2004: SPE+Met

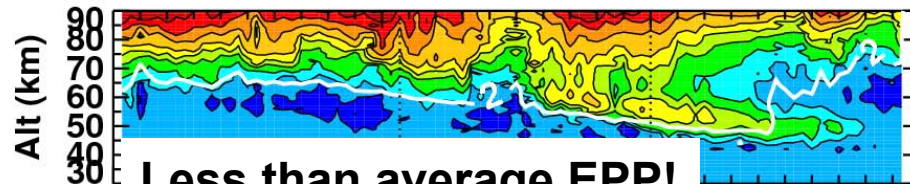
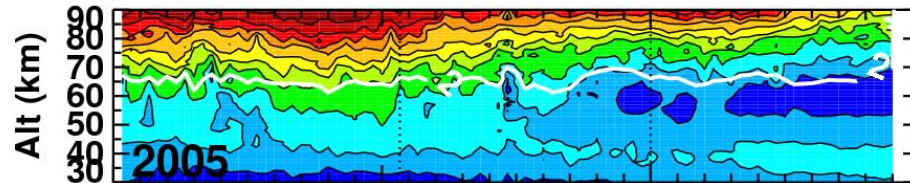
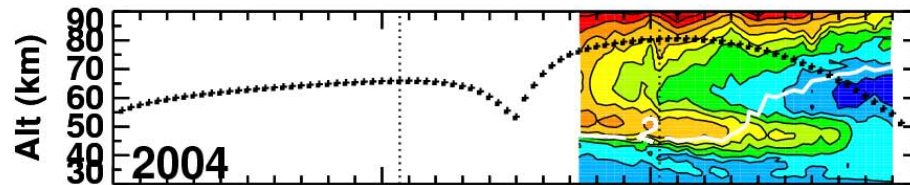
2005-2006: Met



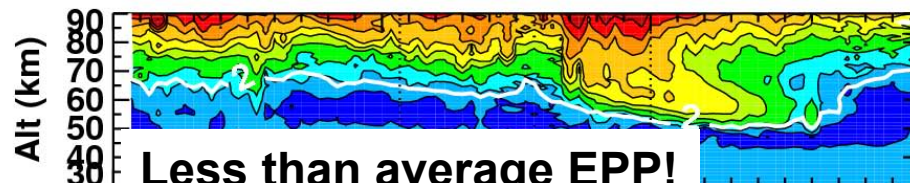
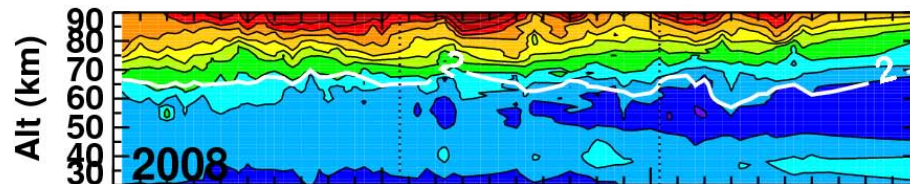
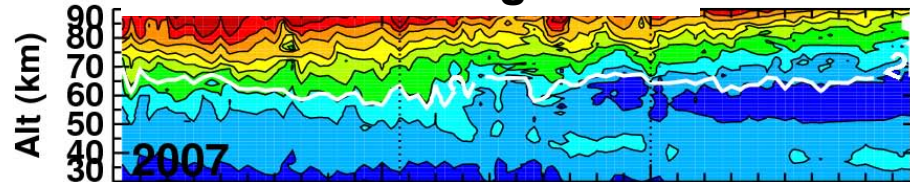
NORTH

ACE NO_x 2004 – 2009

SOUTH



Less than average EPP!

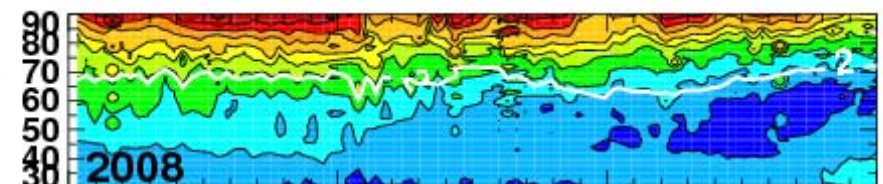
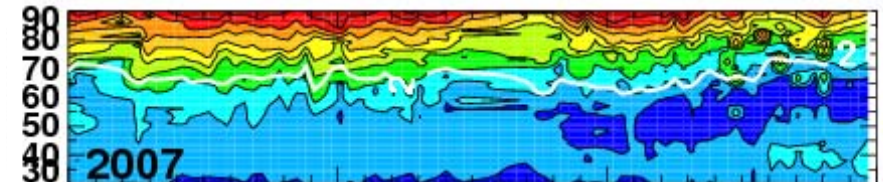
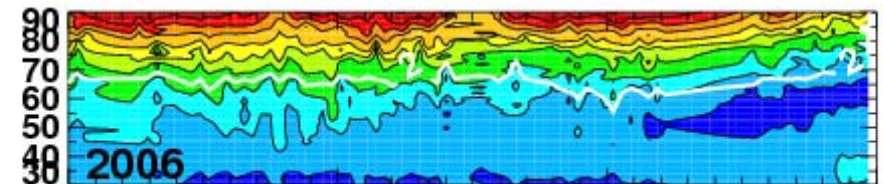
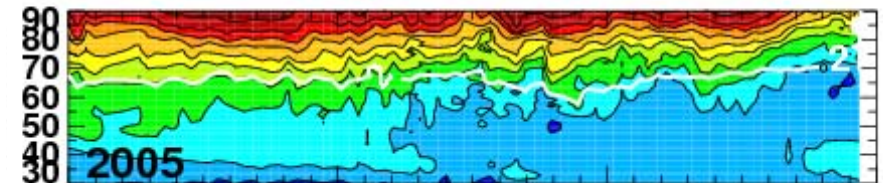
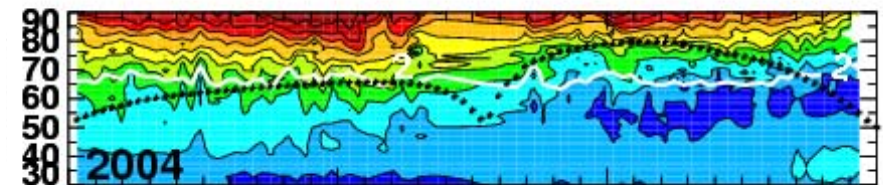


Less than average EPP!

Jan

Feb

Mar



Jul

Aug

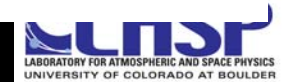
Sep



Randall et al., 2009



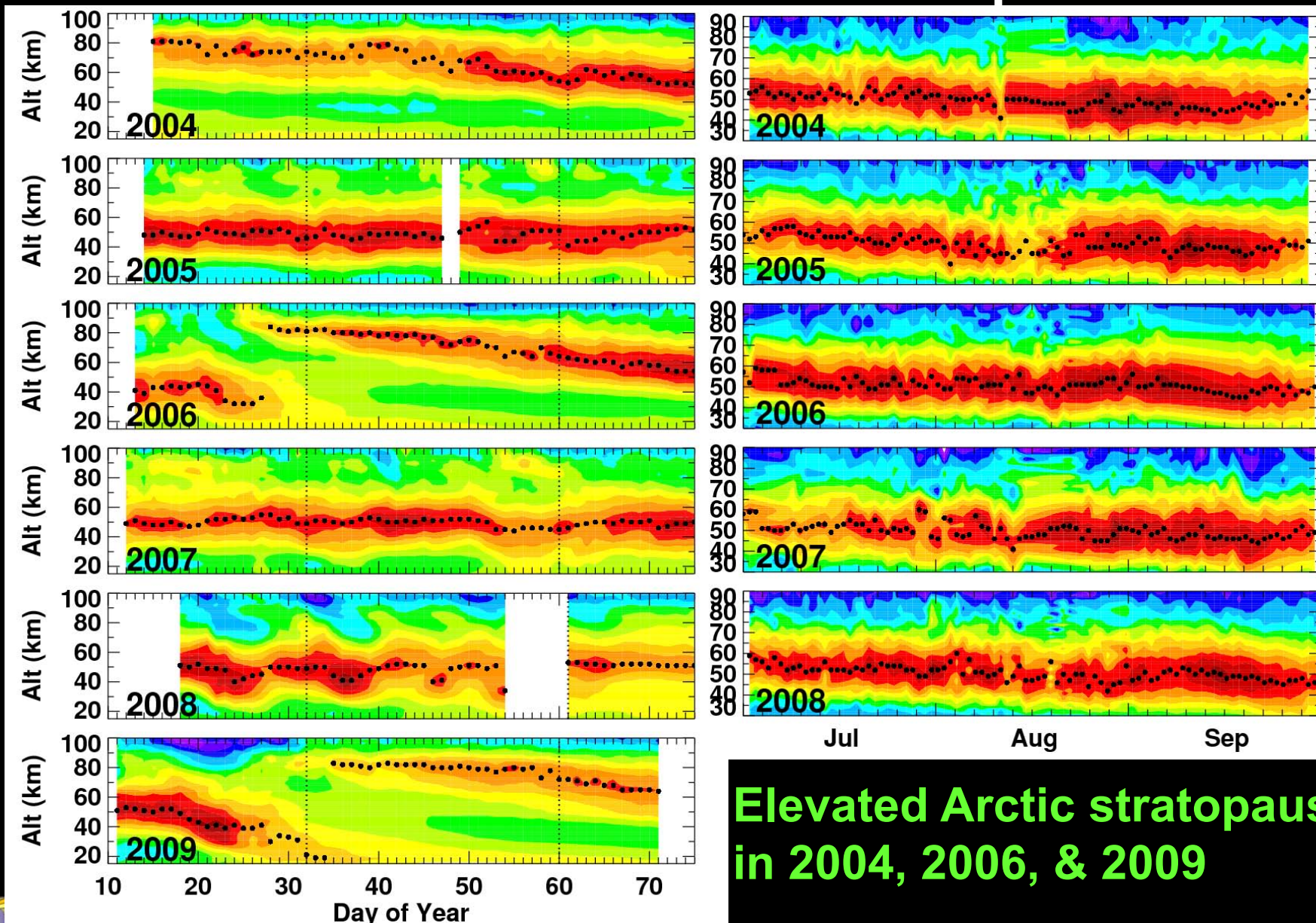
Cora Randall, Aspen Global Change Institute, 14 June 2010



NORTH

SABER & ACE Polar Temp

SOUTH

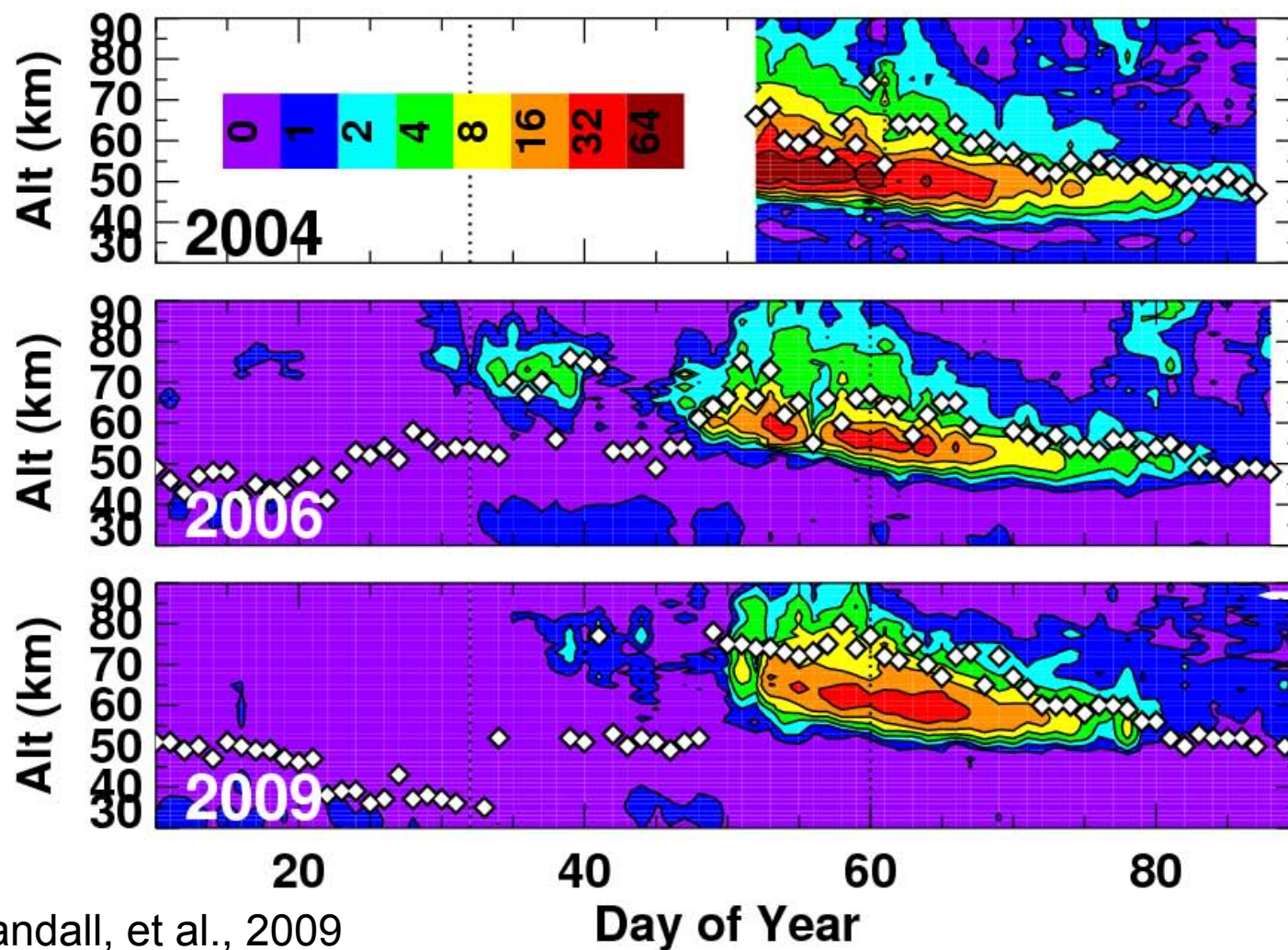


**Elevated Arctic stratopause
in 2004, 2006, & 2009**

Slightly different temperature color scales; range from ~170K-270K (blue-red).



Ratio of observed NO_x in 2004, 2006, & 2009 to average NO_x in 2005, 2007, & 2008



Randall, et al., 2009

NO_x enhancements coincide with elevated stratopause

Elevated stratopause indicative of enhanced descent in (normal) mesosphere region (adiabatic compression).

Enhanced descent brings down more EPP-NO_x.

More efficient transport leads to large stratospheric EPP IE even when EPP itself is well below average.

What is the mechanism?

Begins with extraordinary sudden stratospheric warming (strong and persistent; Manney et al., 2005; 2008; 2009)



- 1. Stratospheric Warming: Equator-to-pole T gradient reverses → Zonal wind reverses direction**
- 2. Planetary waves cannot propagate upward so upper stratosphere & lower mesosphere cool**
- 3. Causes reformation of very strong upper vortex and westerly winds**
- 4. Gravity waves with westward phase speed preferentially propagate to the mesosphere**
- 5. Mesospheric westerly zonal wind slows → induces poleward meridional wind to balance pressure gradient & Coriolis**
- 6. Leads to enhanced descent in the polar mesosphere**
 - Brings down more NO_x**
 - Adiabatic compression results in elevated stratopause**

[Hauchecorne et al., 2007; Siskind et al., 2007; Sahishkumar & Sridharan, 2009; Winick et al., 2009]

Why did we get such impressive stratospheric warmings and recovery in 2004, 2006, & 2009 ?

We don't know.

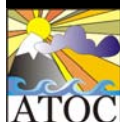
Maybe climate change, but still speculative

- Although 2004, 2006, and 2009 were highly unusual, the frequency of SSW seems to be increasing (~1 per year since 1999, twice as frequent as in previous half-century)
- Models predict more variability with climate change, so extremes might occur more often
- Models inconclusive with regard to frequency & strength of SSWs



The EPP IE Story from Observations

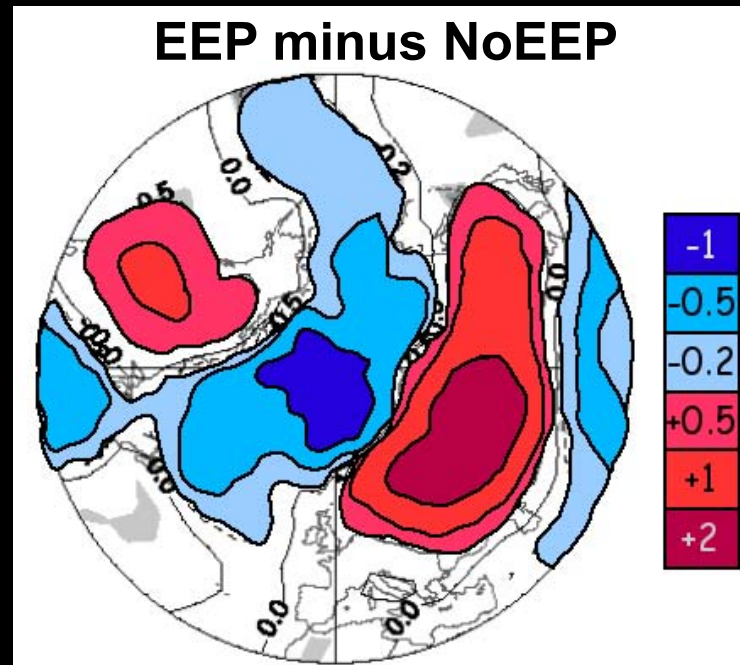
- EPP-NO_x is produced continually and can contribute up to 40% of polar stratospheric NO_x budget even in years with low geomagnetic activity.
- Ozone is depleted by EPP-NO_x by 35% or more.
- Contribution of EPP-NO_x to the stratosphere does not correlate well with the solar (F10.7) cycle.
- Understanding wave mean flow interactions is required to elucidate mechanisms controlling interannual variability.
- Improved picture of EPP effects requires continuous nighttime observations of NO_x throughout the MLT.



Aside #1

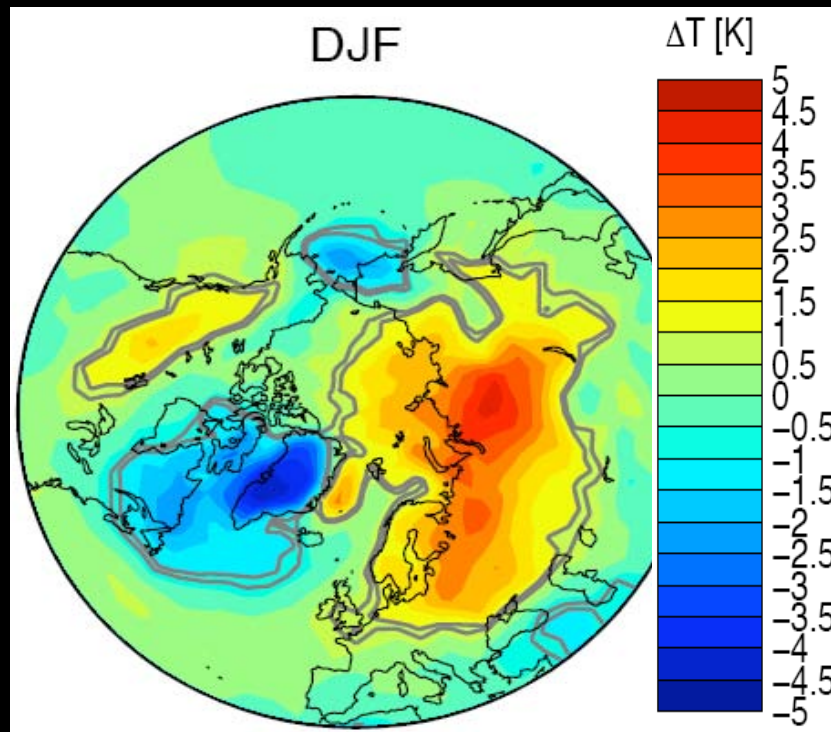
Does EPP affect surface air temperature?

CCM MODEL

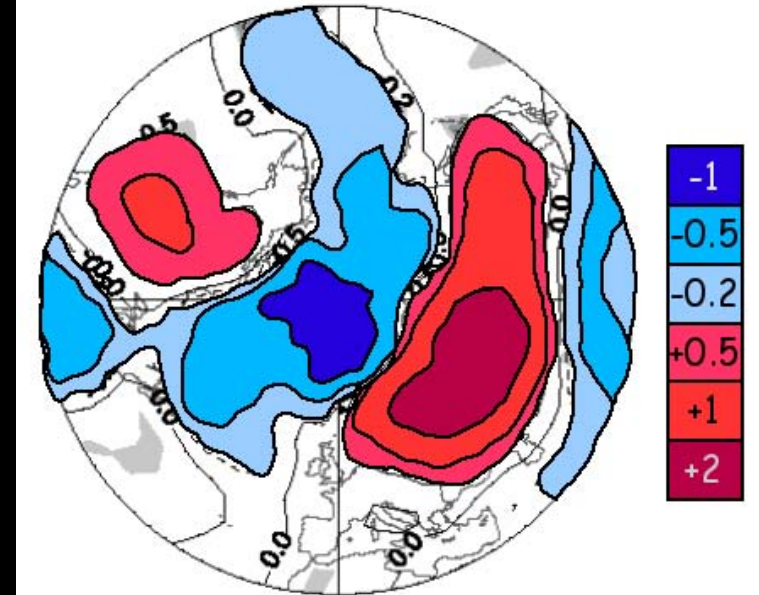


Rozanov et al., 2005. Model results for DJF surface air temperature differences (NH).

Is there empirical evidence that EPP-NO_x affects climate?



Adapted from Rozanov et al., 2005

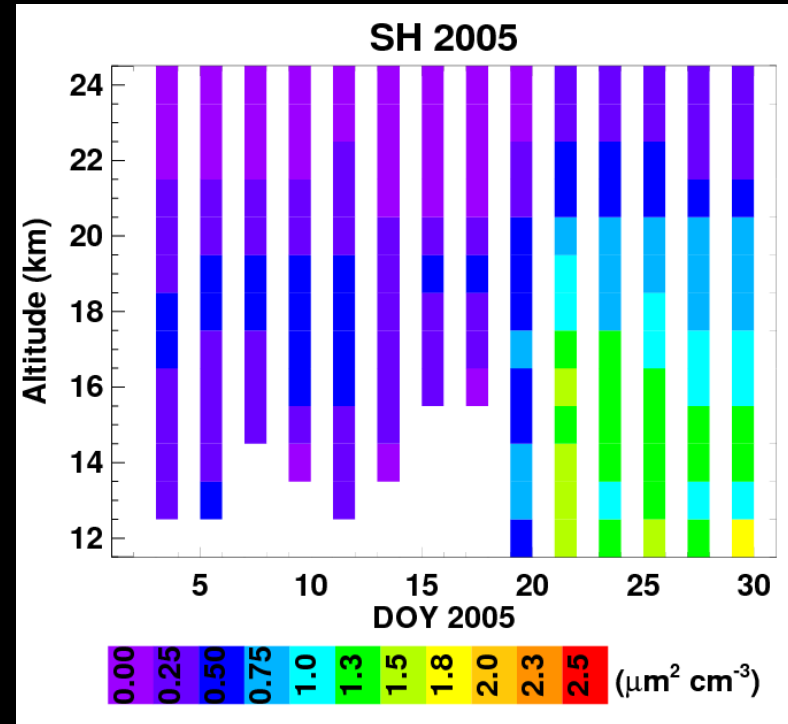
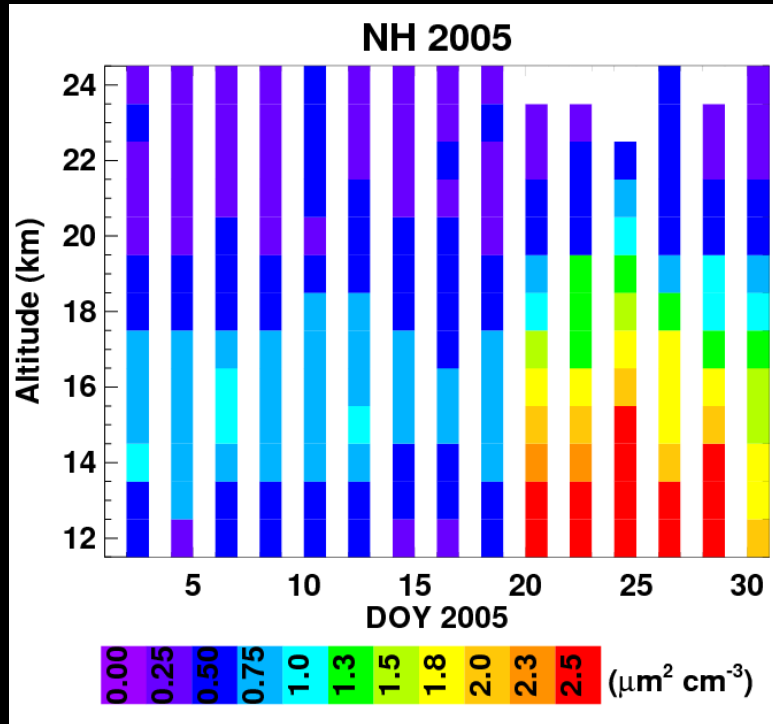


Seppälä et al., JGR 2009:

- Analysis of ERA-40 temperatures yields statistically significant surface air temperature differences between years with high and low EPP.
- Can rule out correlation with solar flux, QBO, ENSO, SAM
- Possible correlation with NAM or random variations in SST

Aside #2

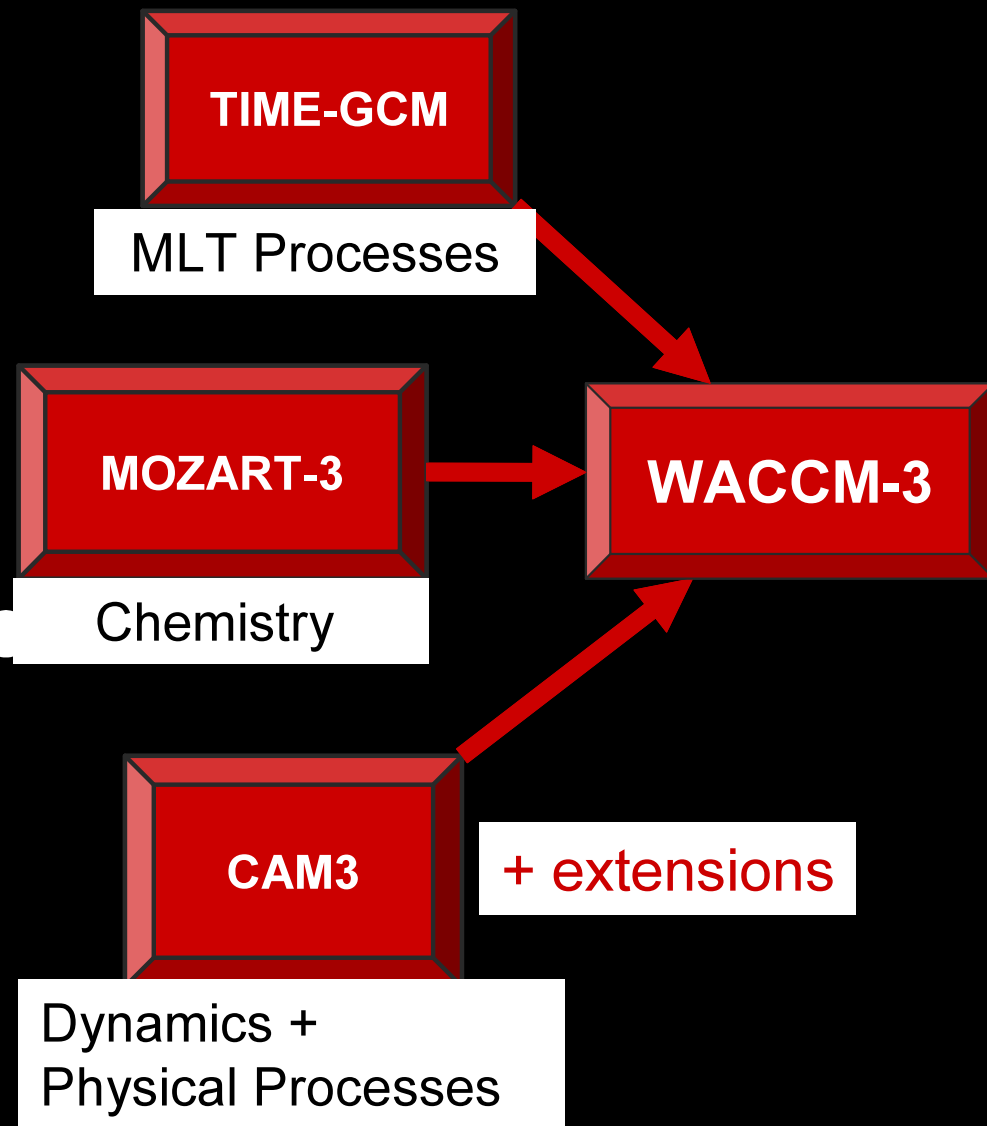
What about EPP effects on clouds/aerosols?



POAM III aerosol surface area increased significantly ($>2\sigma$) after January 2005 SEPs.

- Dynamical cause ruled out.
- No other such increase in 8-year record.
- No information from POAM at lower altitudes

Mirinova et al., in preparation



Whole Atmosphere Community Climate Model

A 3D coupled chemistry climate model

- 0 to ~145 km
- Comprehensive chemistry incl. heterogeneous rx
- Interactive Chemistry or Specified Meteorology
- 1-1.5 km vertical resolution in stratosphere
- 1.9° x 2.5° or 4 x 5° horizontal resolution
- Ref: Garcia et al., 2007

(Adapted from Rolando Garcia)



WACCM Parameterization of Precipitation Effects

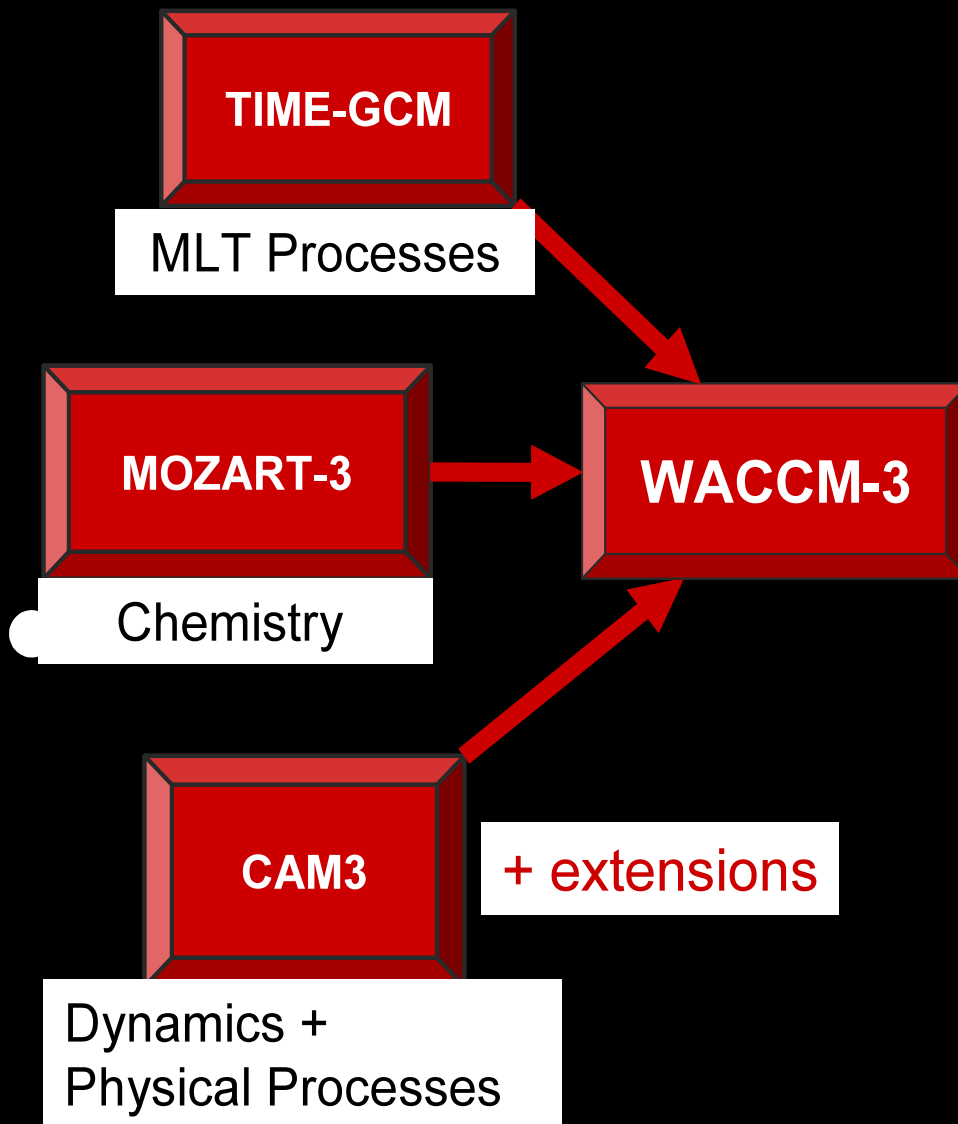
Aurora

- Input = Kp
- Distribution = Auroral Oval
- Roble and Ridley, 1987

Medium Energy Electrons

(30 keV – 2.5 MeV)

- Input = MEPED activity level (or hemispheric power)
- Distribution = Statistical patterns [Codrescu et al., JGR, 1997)
- Fang et al., 2008



(Adapted from Rolando Garcia)



WACCM Runs (free-running coupled climate-chemistry model):

- Seasonally varying (but annually invariant) SSTs
- Constant F10.7 = 210
- 1995 greenhouse gas and halogen concentrations

Daily (repetitive) forcing with 3 cases of EPP:

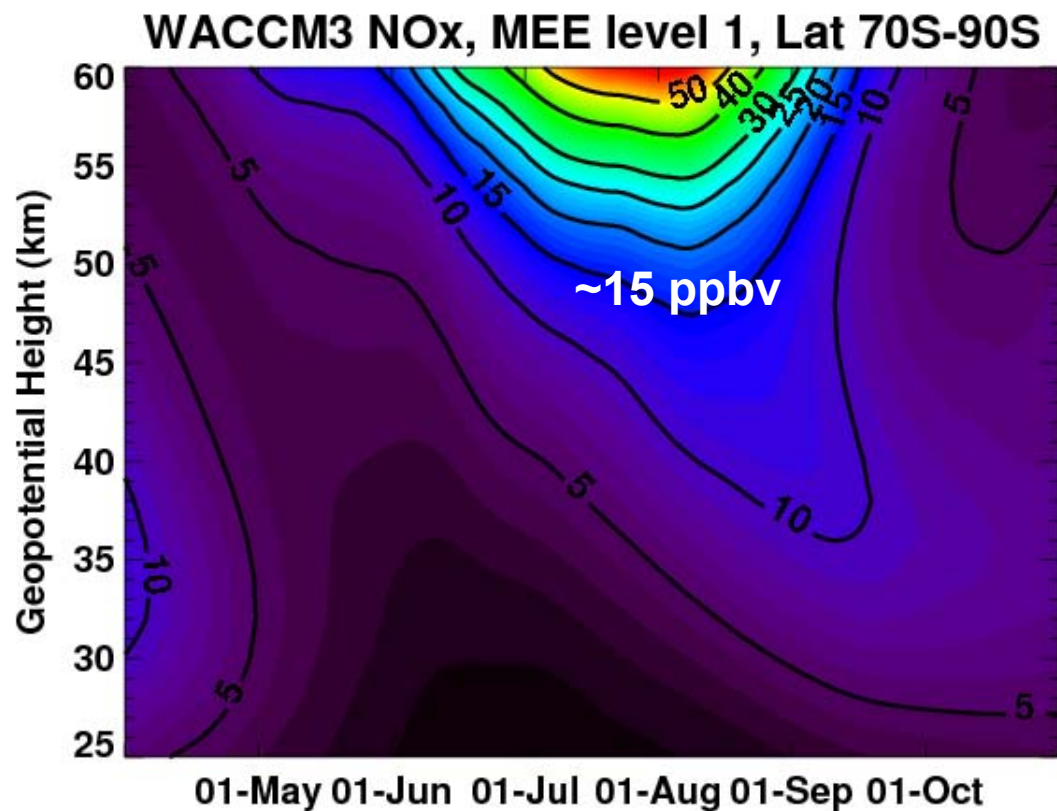
Case 1: ~Zero particle precipitation, $K_p=2/3$ ($A_p=3$)
80 years (after spin-up) = 80-member ensemble

Case 2: Moderate auroral electrons, $K_p=4$ ($A_p=27$)
50 years (after spin-up) = 50-member ensemble

Case 3: Moderate auroral electrons plus >30 keV electrons (level 1; occurs 40% of the time)
50 years (after spin-up) = 50-member ensemble



WACCM simulation similar to MIPAS



- 70-90° S Latitude
- Case 3
(realistic simulation of 2003 geomagnetic activity)

But WACCM
underestimates EPP-NO_x
by about a factor of 2.

50 km

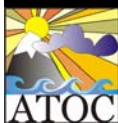
50 km

40 km

40 km

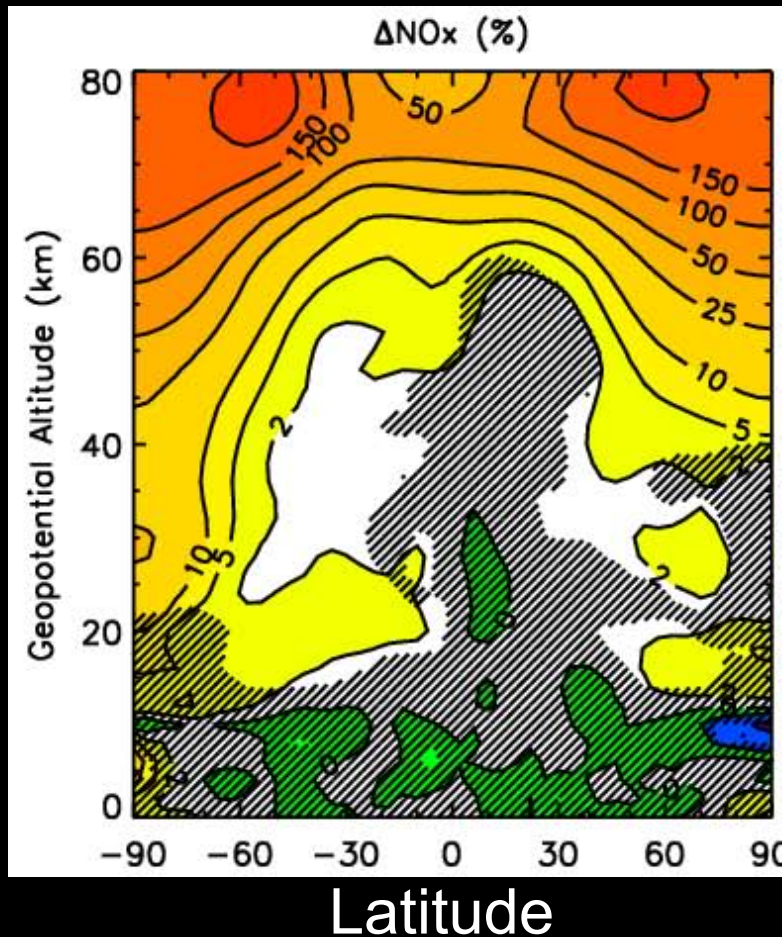
30 km

30 km

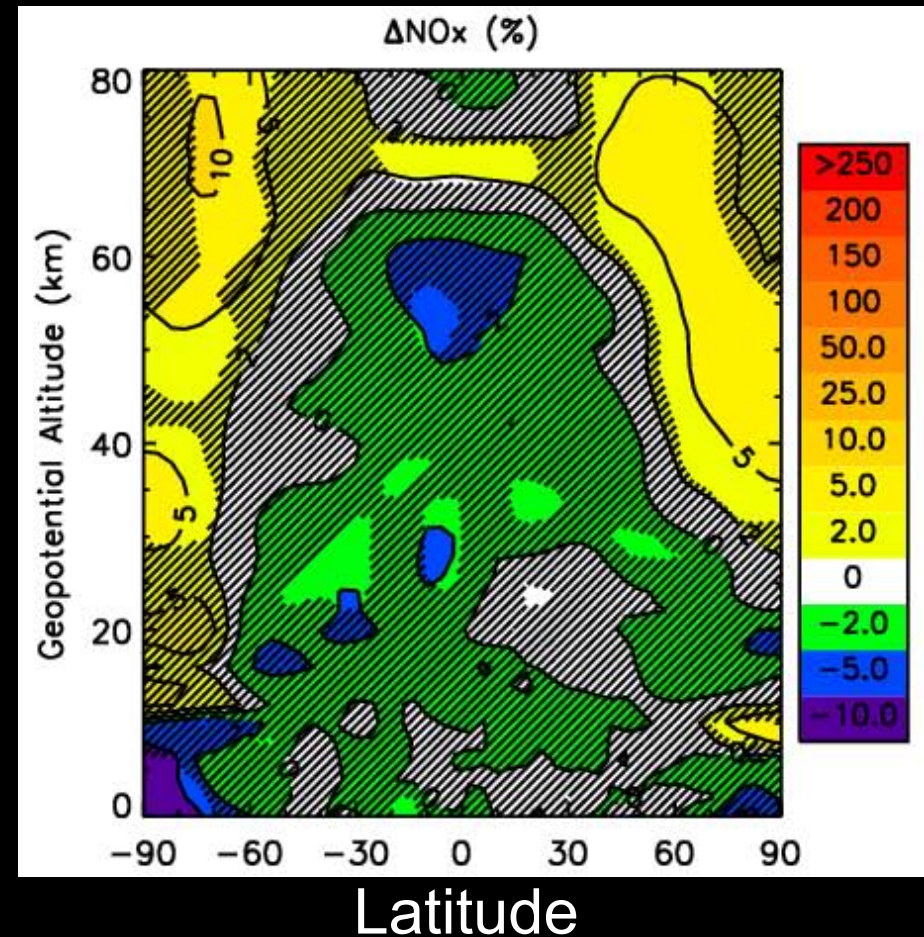


Change in NO_x due to EPP: Annual Averages

Aurora Effect Case 2 minus Case 1



MEE Effect Case 3 minus Case 2



Regions without cross-hatching significant at 95% level

Cora Randall, Aspen Global Change Institute, 14 June 2010

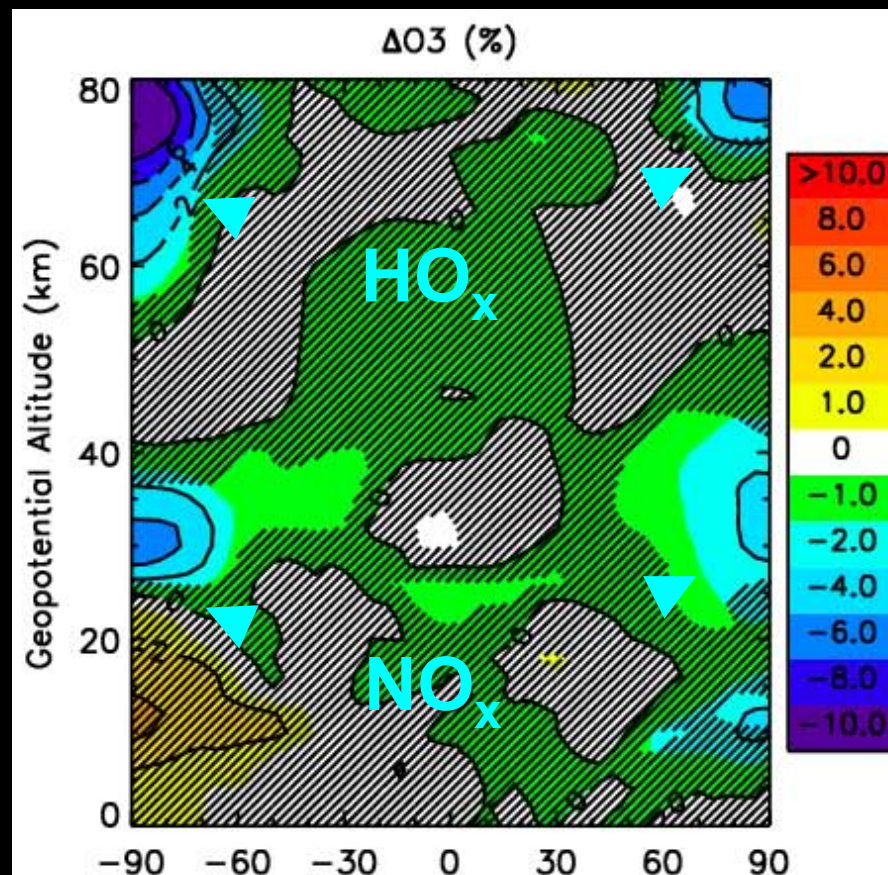
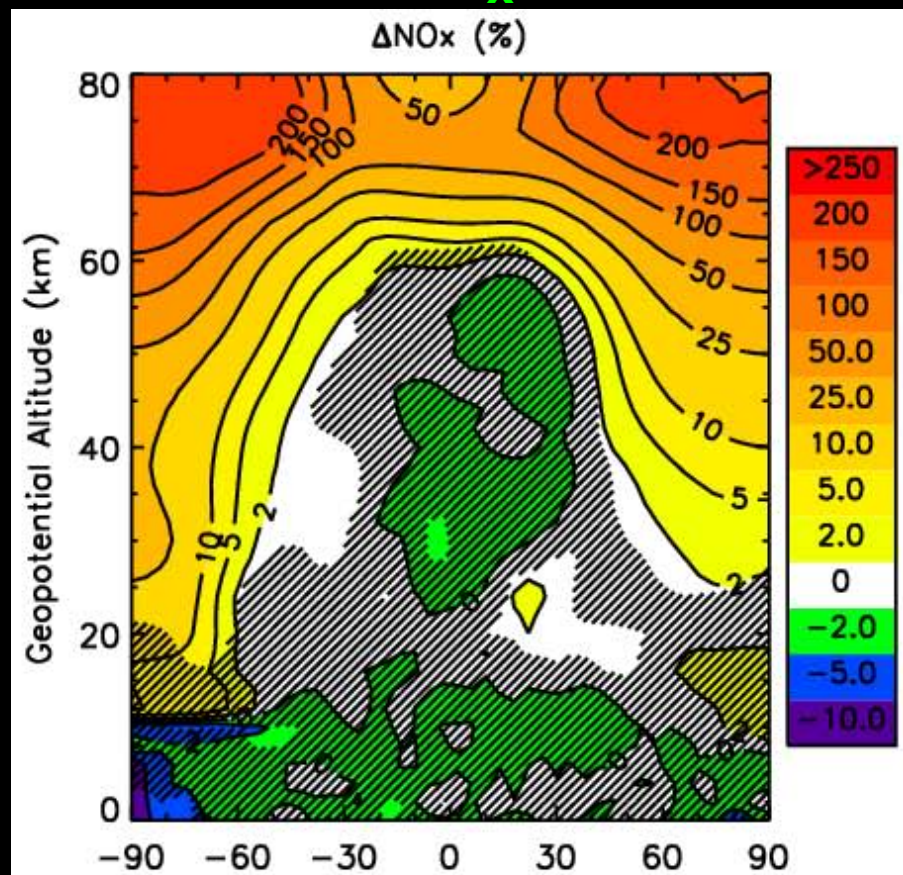


Change in NO_x and O_3 due to EPP

Annual Averages: Aurora + MEE (Case 3 minus Case 1)

NO_x

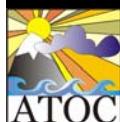
Ozone



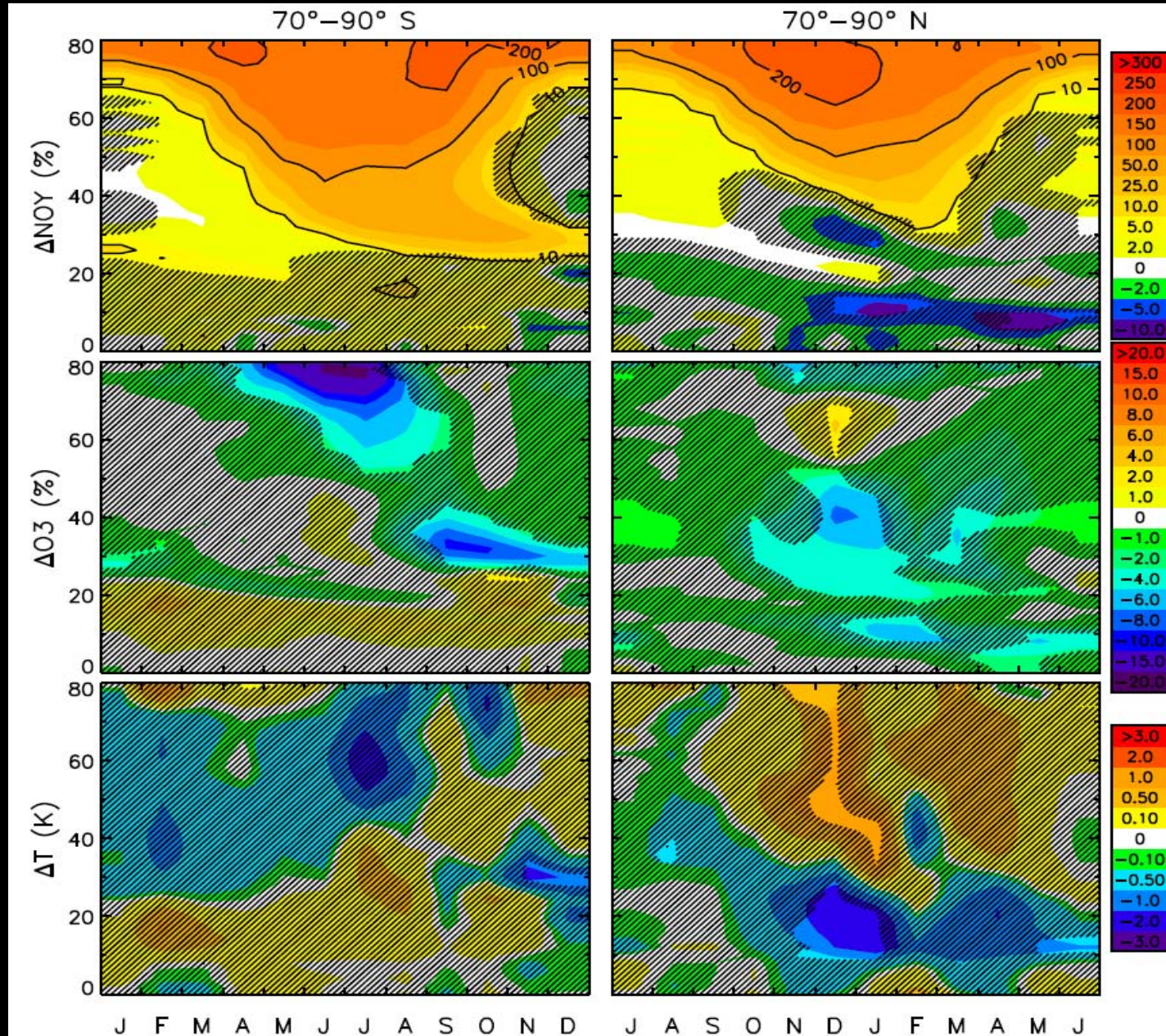
Latitude

Latitude

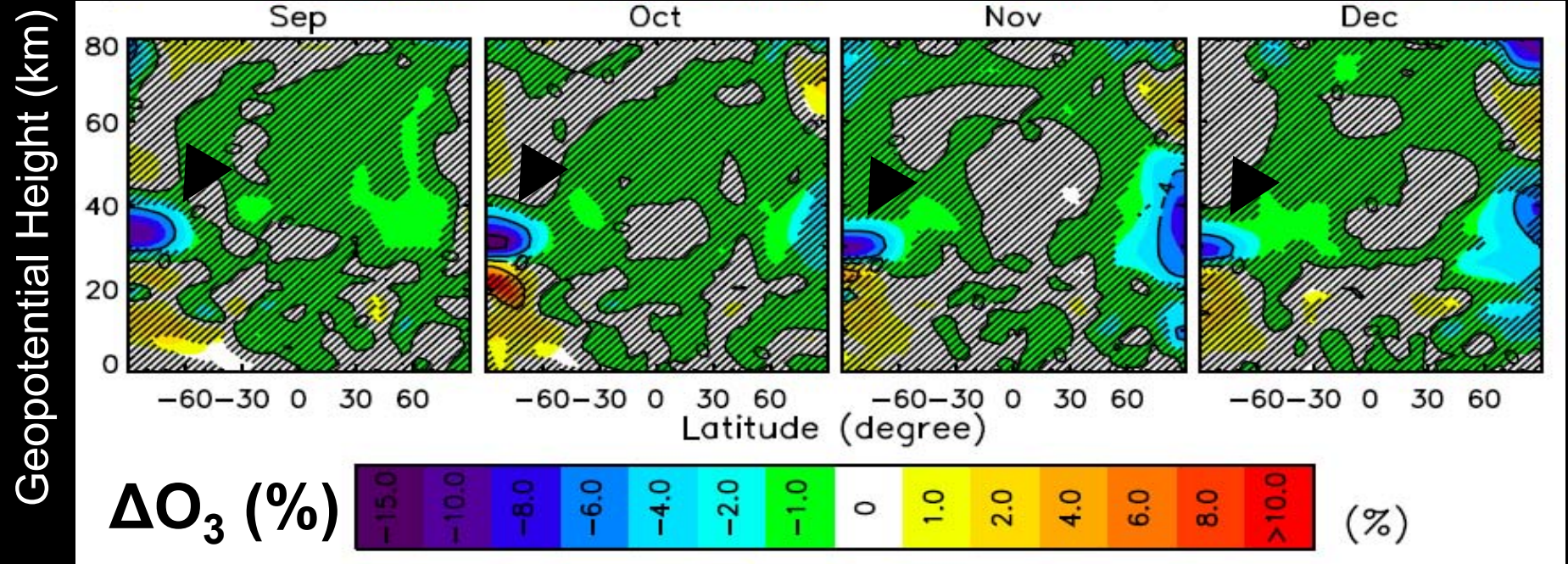
Regions without cross-hatching significant at 95% level



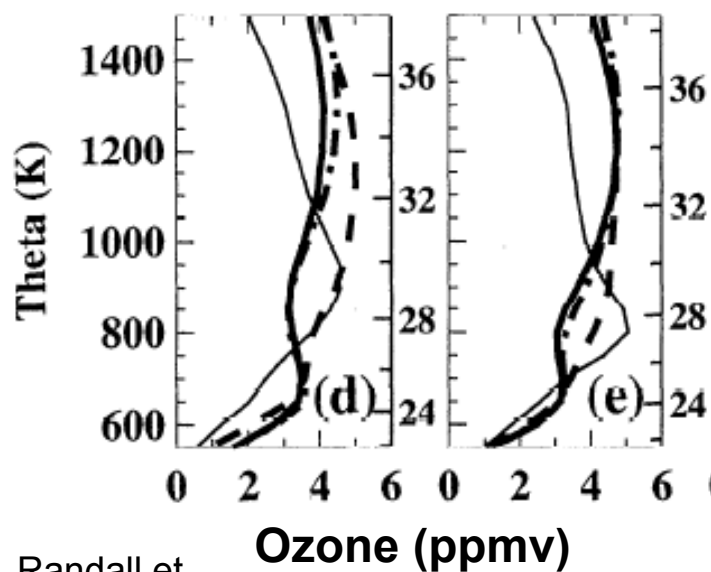
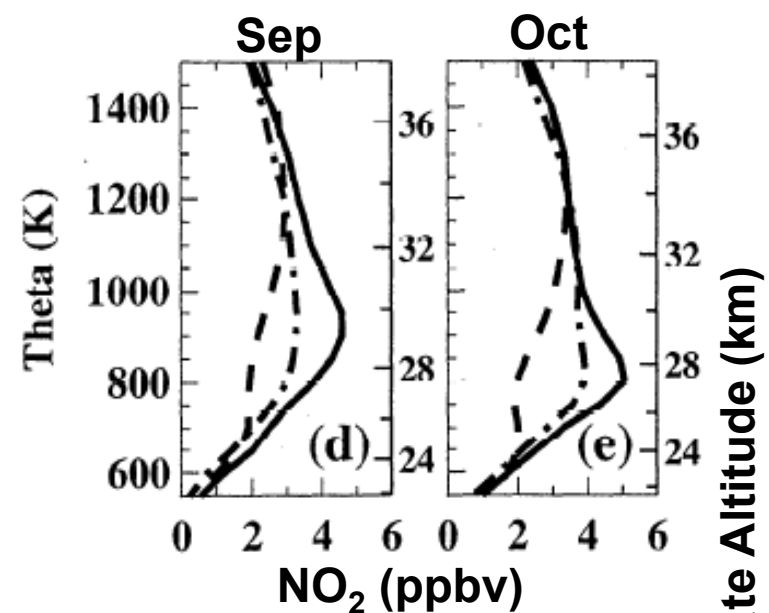
Auroral Effect Time Series (Case 2 minus Case 1)



Monthly average ozone depletion of up to 15% at high southern latitudes, 30-40 km

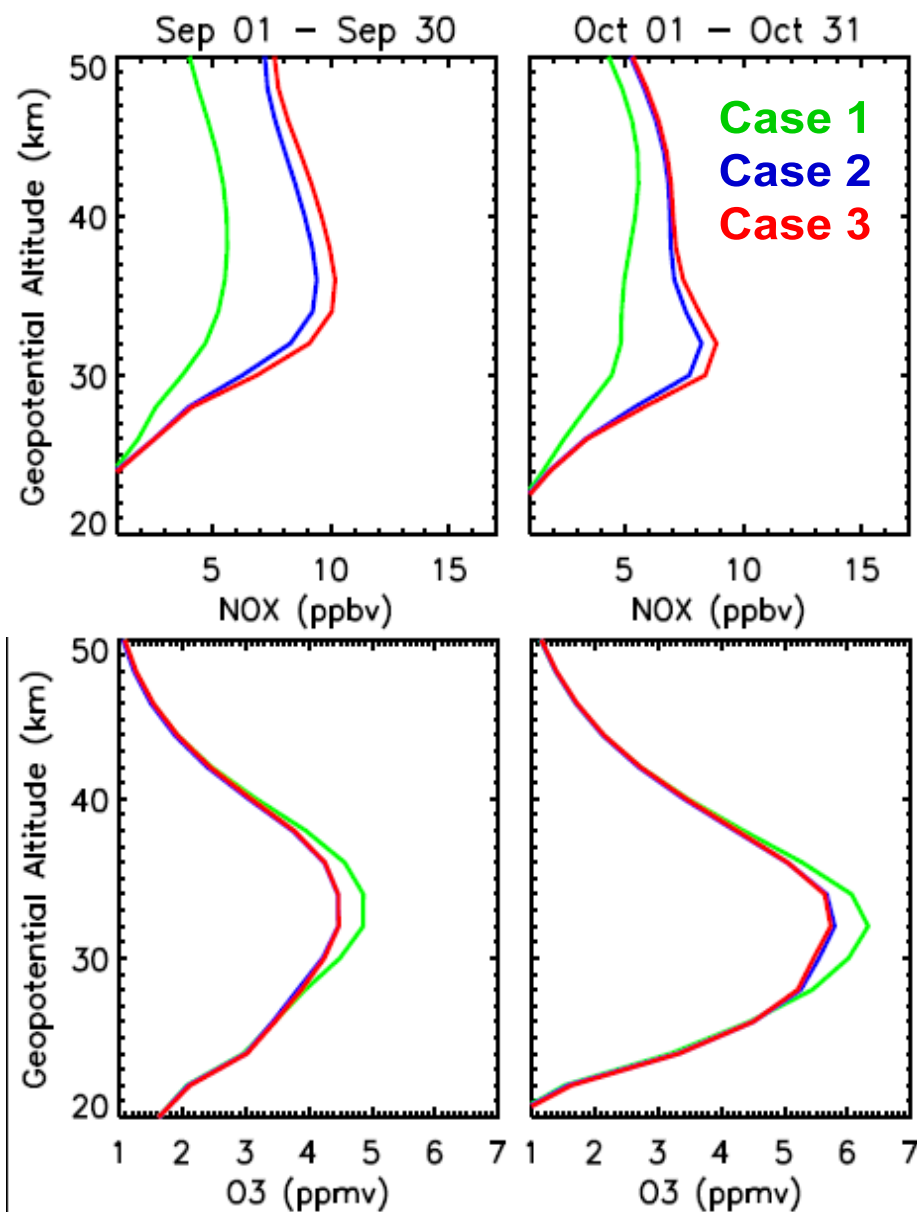


Corresponds to catalytic NO_x destruction



Randall et al., 1998

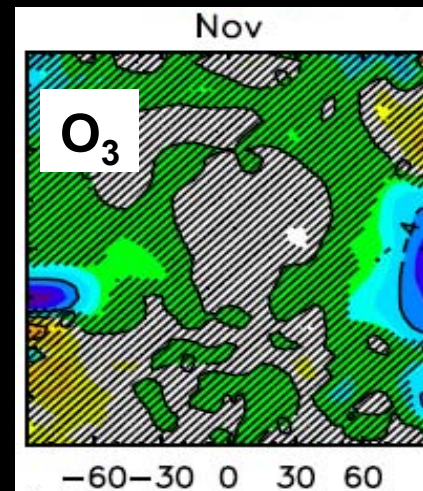
POAM
Data 1994
1996



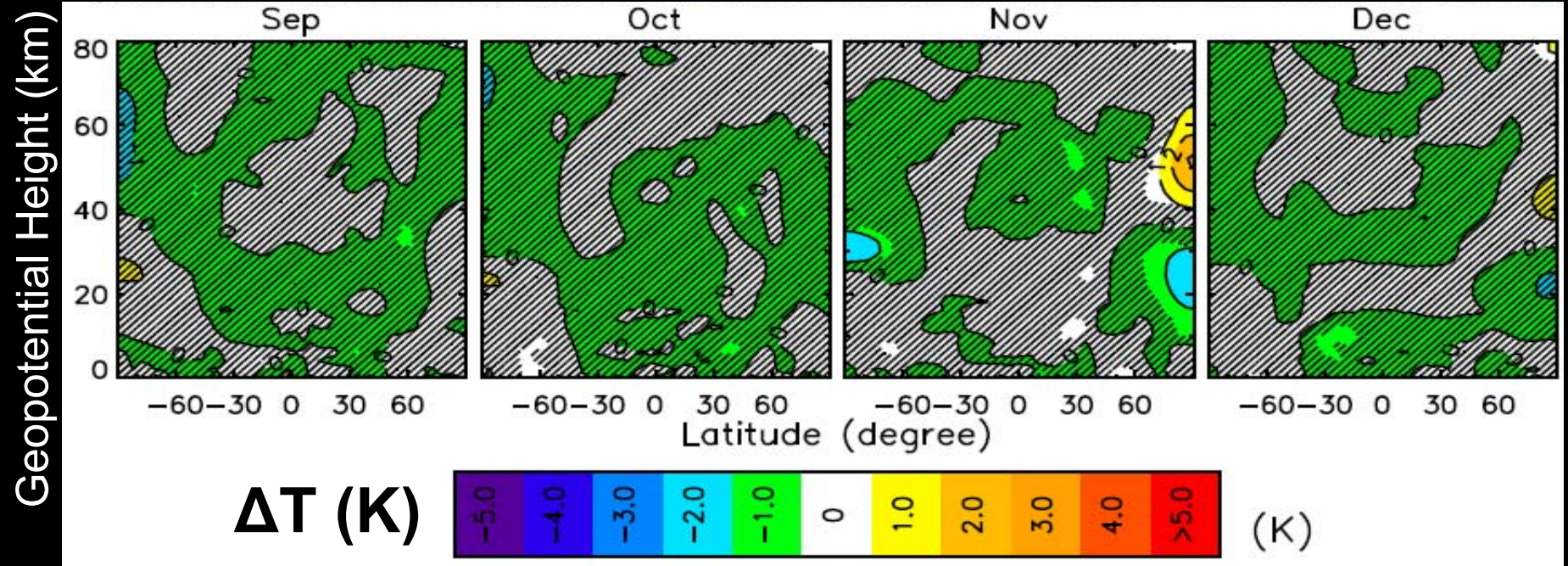
WACCM

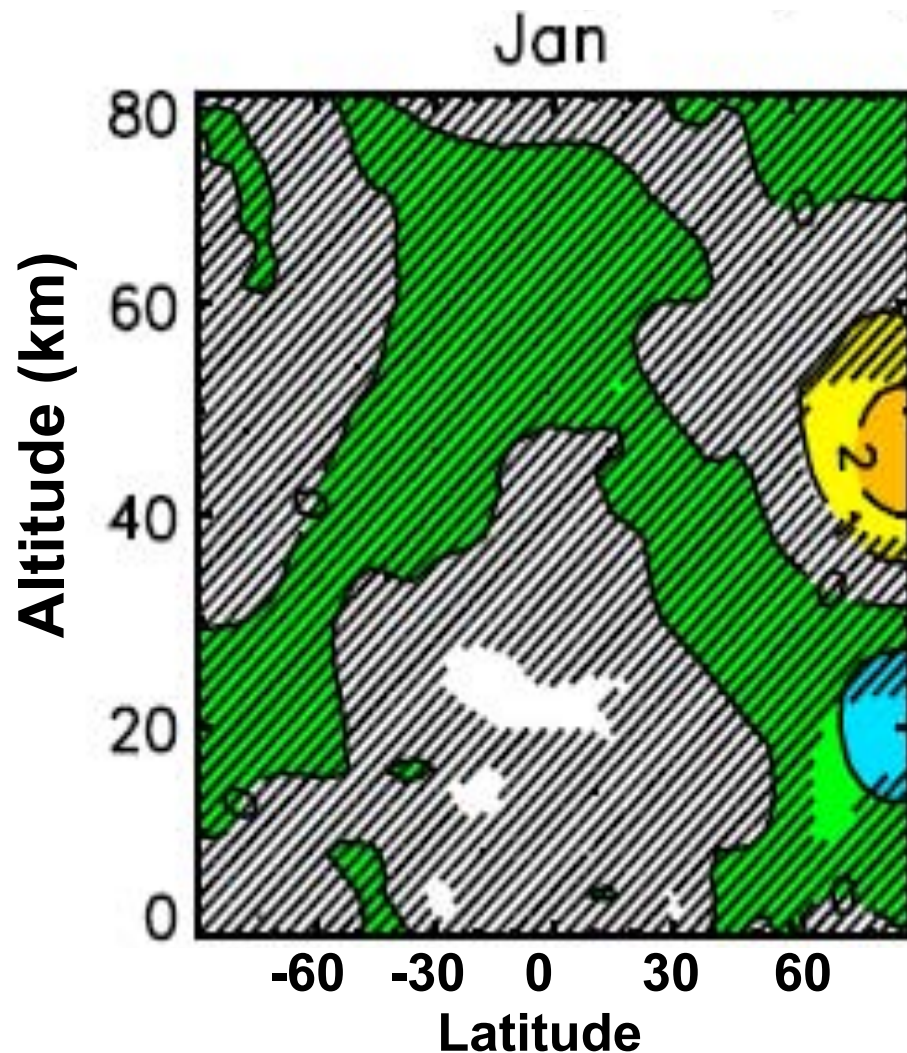
Up to 2 K monthly average temperature changes due to Aurora + MEE.

Are these changes consistent with ozone?



Is heating due to less radiative cooling by O₃ in polar night upper stratosphere?





Are temperature effects valid?

Still trying to understand statistical significance:

Statistically significant differences found between two sets of 80-member ensembles, both with same forcing.

Summary of WACCM Model Results

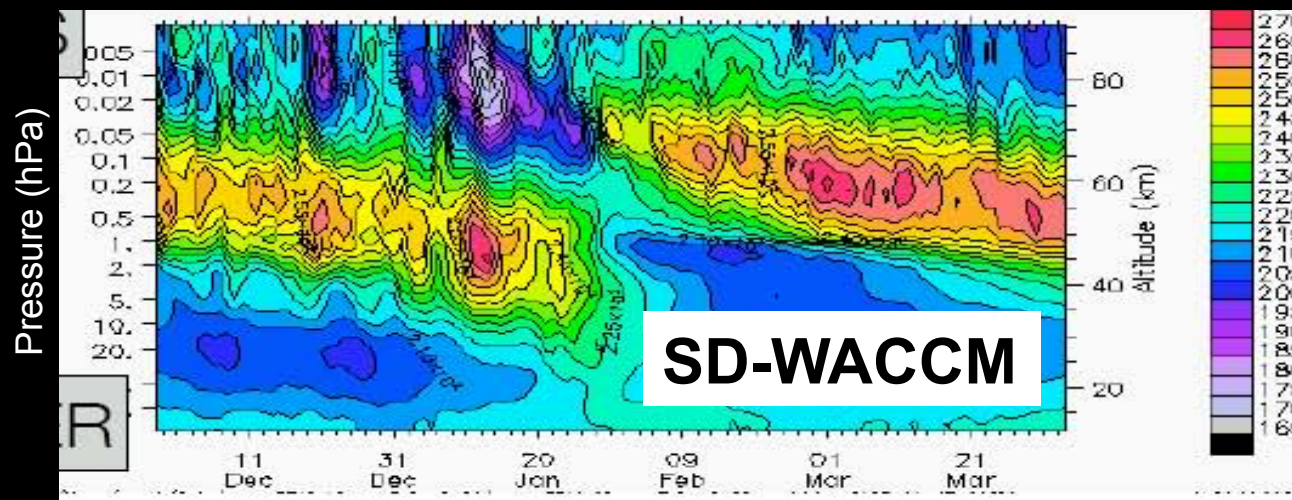
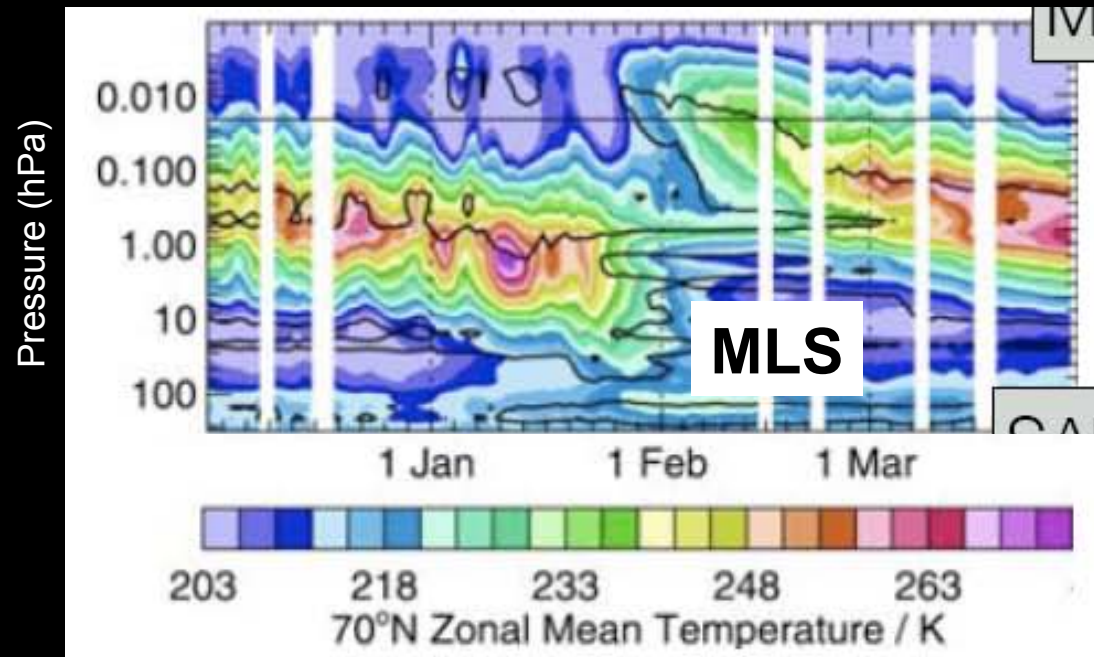
- Free-running WACCM-EEP compares well qualitatively to observations for typical (i.e., low) EEP conditions
 - Clear descent of EPP-NO_x
 - Corresponding ozone depletion
- Descending NO_x underestimated by factor of ~2 even for standard meteorological conditions
- Concerns about T statistics in polar stratosphere
- Specified SSTs, so cannot probe tropospheric effects
- Never see elevated stratopause as in 2004, 2006, 2009



SD-WACCM vs. MLS Temperatures, 2006

If WACCM is nudged
to Met assimilation
(e.g., GEOS-5)
temperature & winds
in stratosphere,
elevated
stratopause
is captured.

Still must
modify
SD-WACCM
for MEE.



Figures courtesy of Dan Marsh, NCAR



Conclusions & Outstanding Questions

EPP DE (Jackman talk) and IE significantly affect polar stratospheric NO_y and Ozone.

Temperature effects not yet quantified.

WACCM underestimates EEP effects for baseline conditions.

Free-running model does not simulate unusual meteorology of 2004, 2006, 2009.



Some Questions

What is NO_x really doing in the polar night?

What are the detailed dynamical mechanisms that link EPP to atmospheric change?

Does EPP affect surface level temperatures?

Are coupling mechanisms from the surface to the MLT changing (more favorable for enhanced EPP- NO_x descent)?

Is there feedback between EPP and the coupling?



Thanks very much!



Cora Randall, Aspen Global Change Institute, 14 June 2010

